

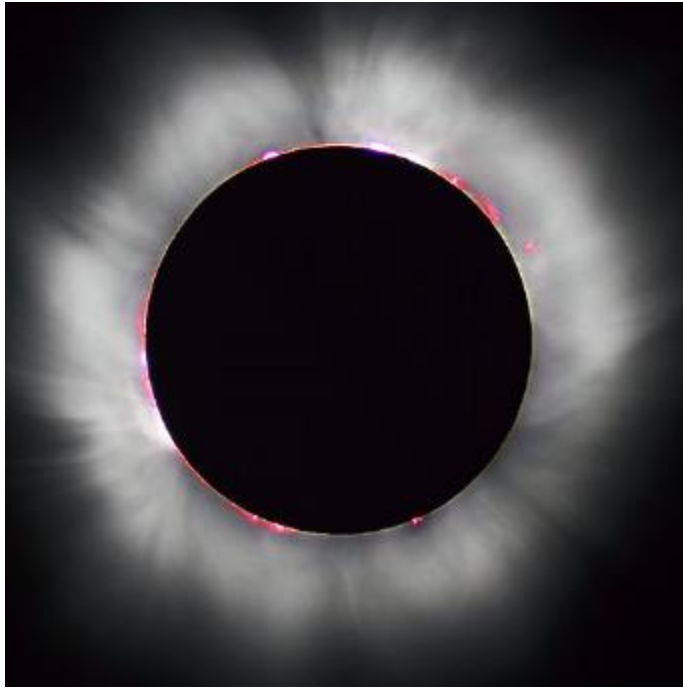
# Eclipses

Interactions of Sun, Earth, & Moon

# the Visible Sky

- I. Stars and Celestial Sphere  
Constellations & Coordinates
- II. Sun  
Time, Seasons, Precession
- III. Moon  
Phase, Orbit, etc.
- IV. Eclipses**  
**Solar & Lunar**

The student will be able to:		HW:
1	Explain and utilize constellations and asterisms as means of mapping and organizing the stars.	1 – 4
2	Explain and utilize the concept of the celestial sphere as a means of understanding the appearance of the universe as seen from Earth.	
3	Explain the significance of the pole star, Polaris, and its connection with the apparent motion of the celestial sphere.	
4	Explain, define, and utilize the celestial equatorial coordinate system of right ascension and declination, celestial equator and celestial poles.	
5	Describe changes in position and appearance of the stars through time and explain in terms of the actual motion and position of the Earth.	5
6	Define, apply, and relate to astronomical events or cycles the following time concepts: sidereal and solar day, sidereal and tropical year, mean solar time, standard time, daylight savings time, and universal time.	6
7	Use a planisphere to locate celestial objects for a particular date and time and/or determine the date and time of certain celestial events.	7 – 8
8	Describe changes in position and appearance of the Sun through time and explain in terms of the actual motion and position of the Earth.	9
9	State the constellations of the zodiac in order and explain the relation between the zodiac and the Sun.	10 – 14
10	Explain, define, and utilize the concept of the ecliptic and the ecliptic plane.	
11	Illustrate and describe the connection between the seasons and the motion and orientation of the Earth in its orbit.	15
12	Explain the cause and effect of Earth's precession and state and apply the period of this cycle to solve problems.	16
13	Describe changes in the appearance of the Moon over the course of one day and night, from one night to the next, from one week to the next, from one month to the next, and from year to year.	17 – 20
14	Explain the apparent motion and changing appearance of the Moon in terms of the actual motions of the Earth and Moon relative to the Sun.	
15	Explain and illustrate how the motion and position of the Moon relative to the Earth and the Sun result in the phases: new Moon, waxing crescent, first quarter, waxing gibbous, full Moon, waning gibbous, third quarter, and waning crescent.	
16	Define, apply, and relate to astronomical events or cycles the following concepts: sidereal month, synodic month, lunar sidereal and solar days.	21 – 22
17	Explain and illustrate how the motions and positions of the Earth, the Sun, and the Moon result in lunar and solar eclipses – partial, total, and annular.	23
18	Explain and illustrate the concepts of umbra and penumbra in relation to eclipses.	24



## Solar Eclipse

The Sun is blocked (eclipsed) by the Moon.



## Lunar Eclipse

The Moon is blocked (eclipsed) by the Earth (*i.e.* Earth's *shadow*).

# Solar Eclipses



Partial



Annular

What phase is  
the Moon?

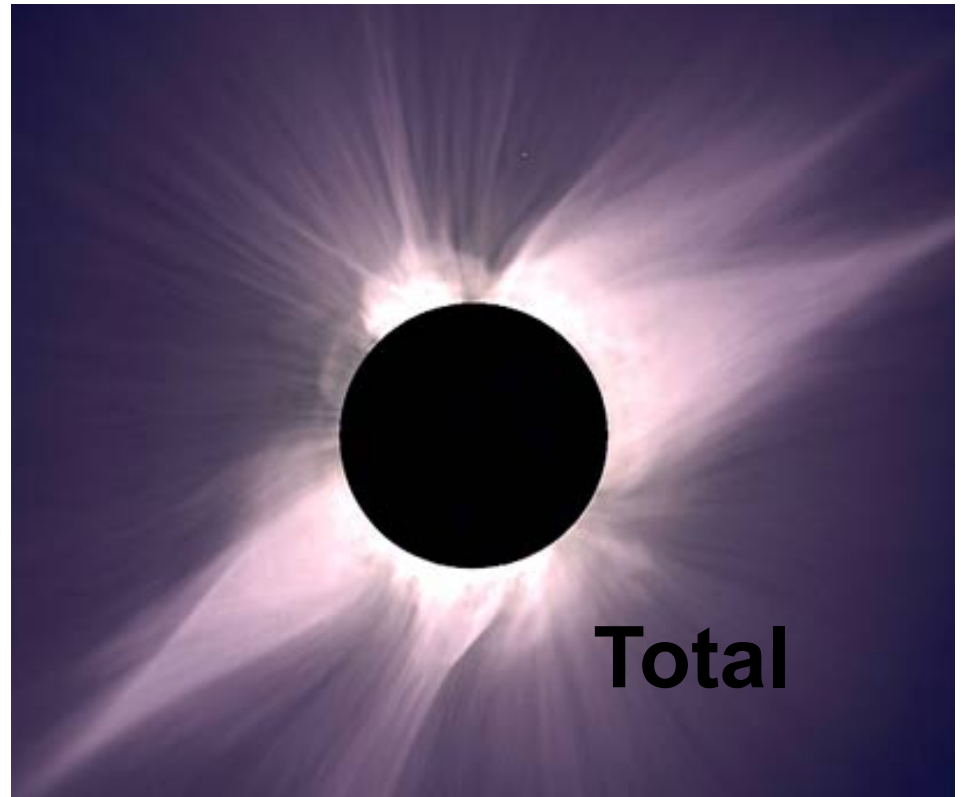
New!

New Moon at *every*  
solar eclipse?

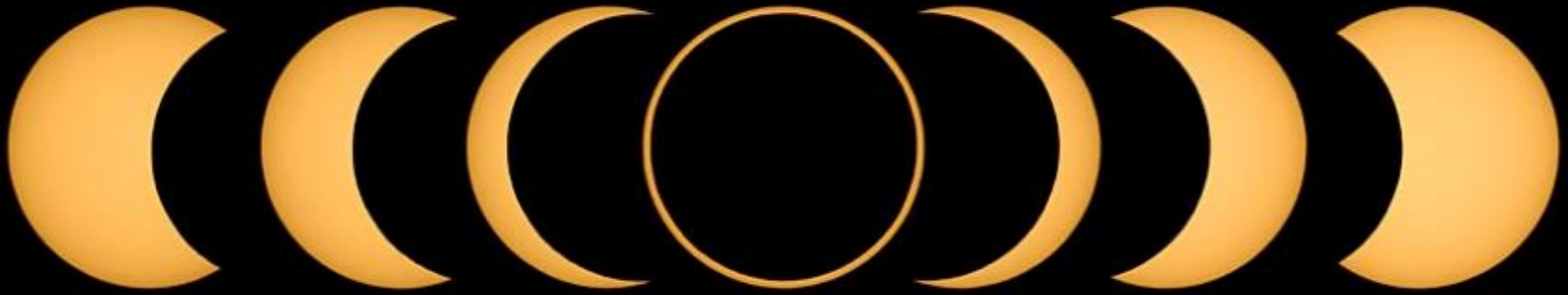
Yes!

Solar eclipse at *every*  
New Moon?

No!



**Total**



[www.MrEclipse.com](http://www.MrEclipse.com)

©2005 F. Espenak

Time lapse sequence of an annular eclipse.

The Moon isn't "big enough" during this type of eclipse to *completely* block out the Sun. Why not?!

This occurs when the Moon is far enough away from Earth and/or the Sun is close enough to Earth. These distances vary because neither orbit is perfectly circular.

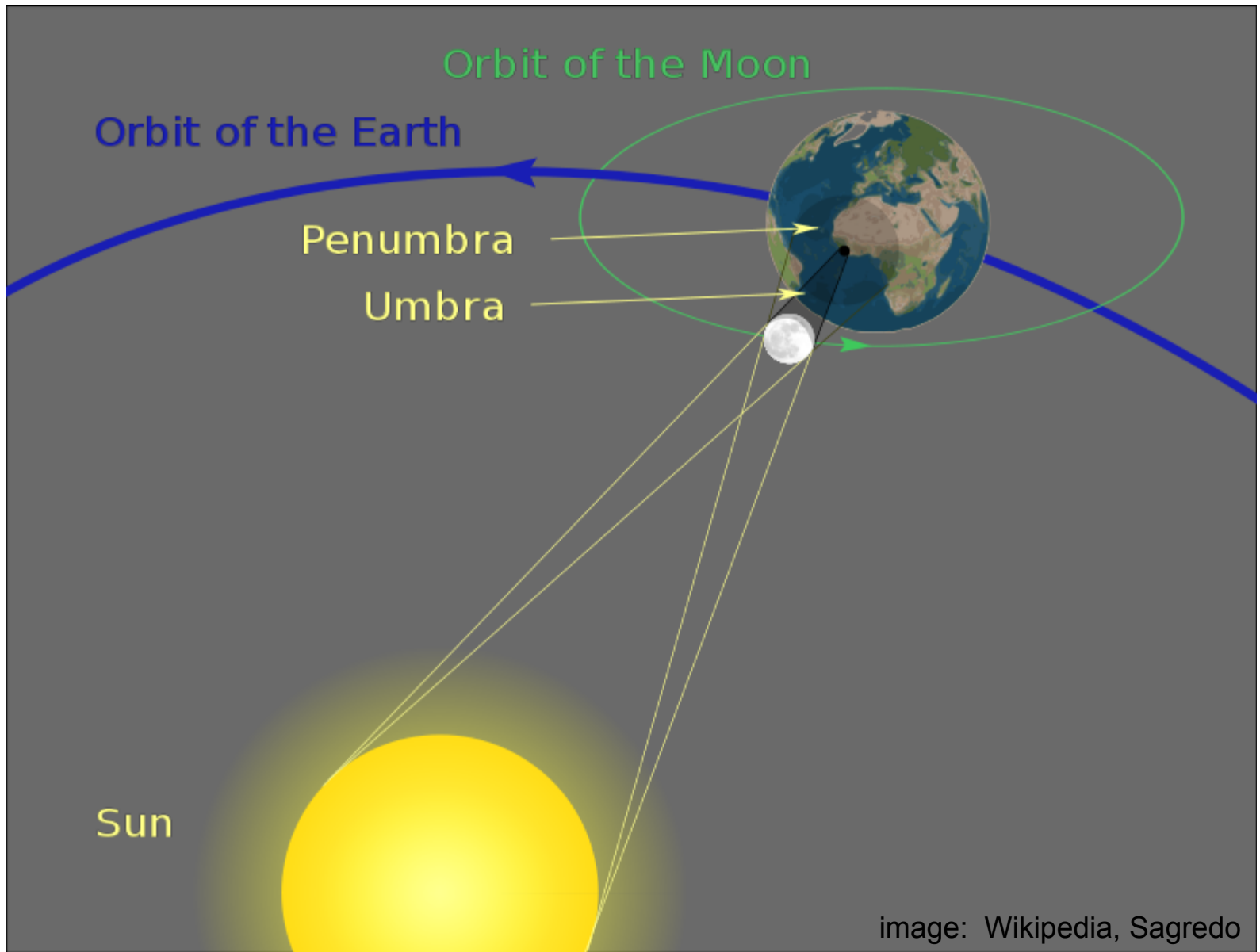
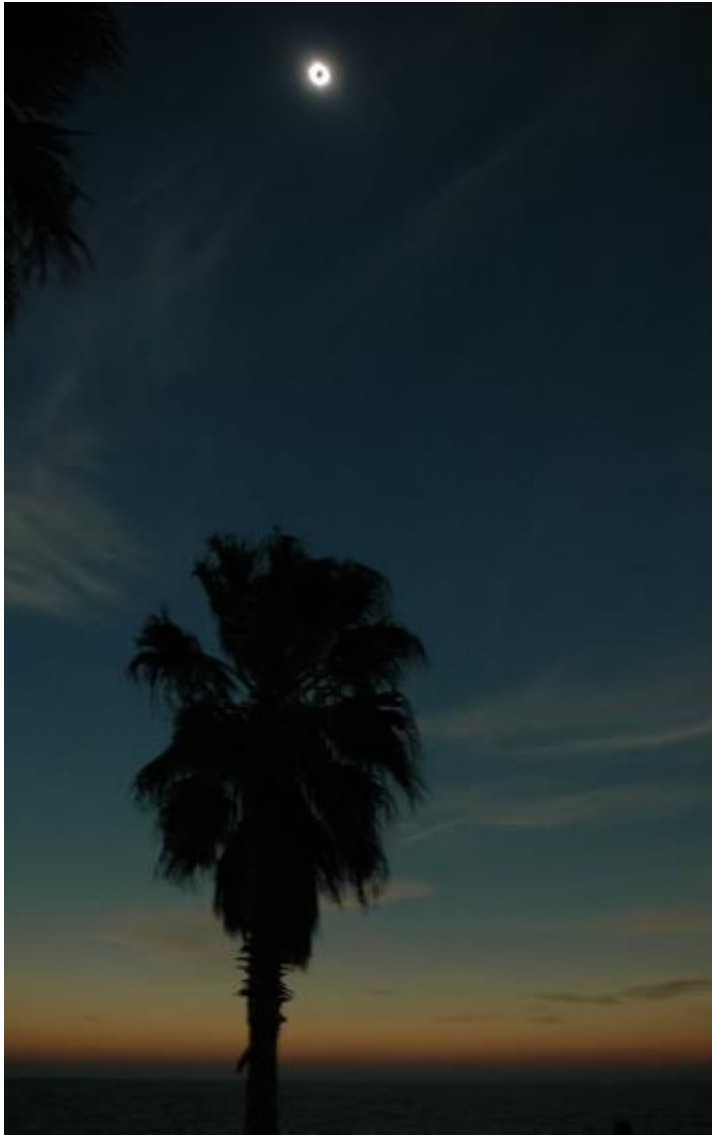


image: Wikipedia, Sagredo

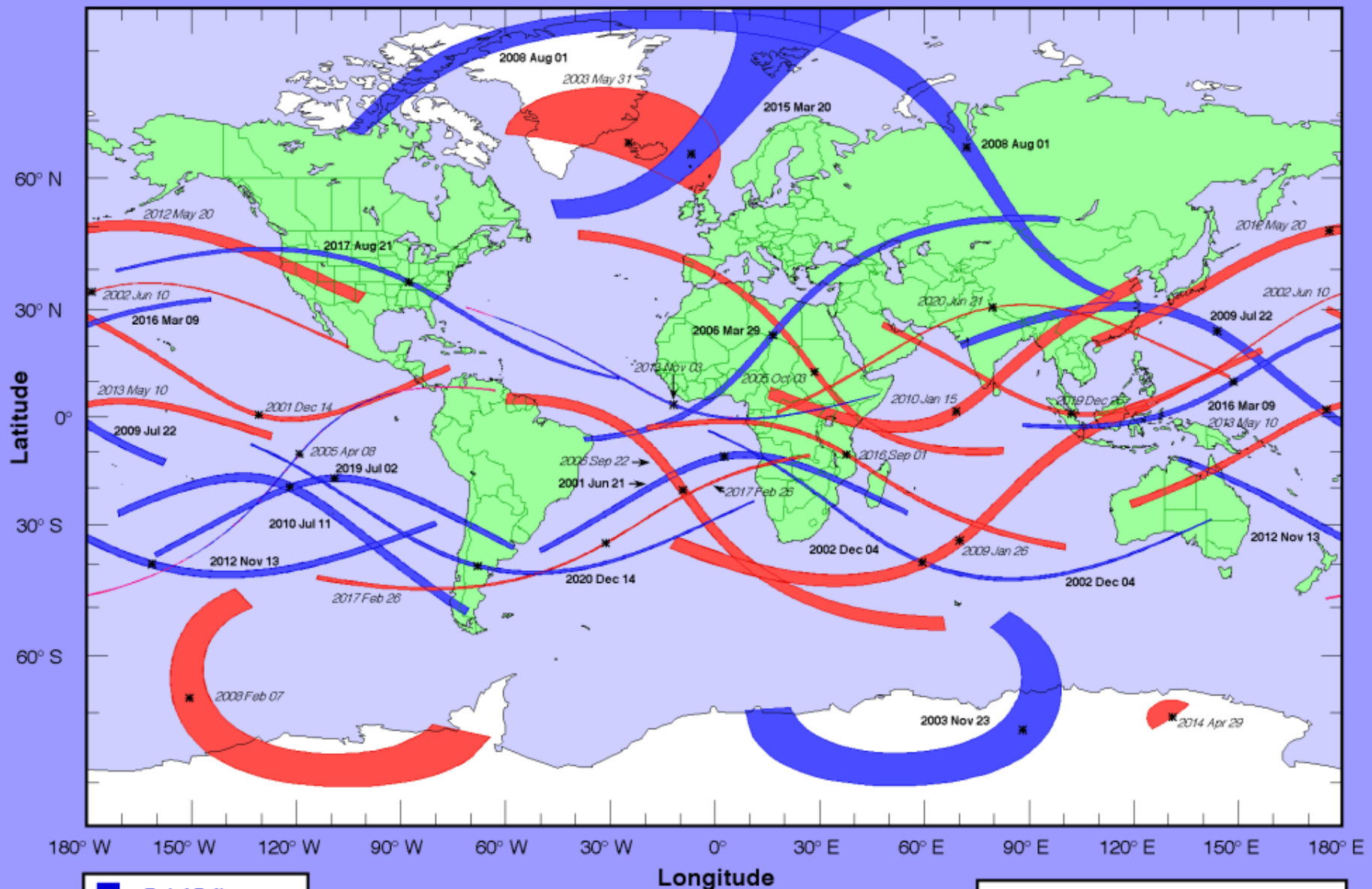




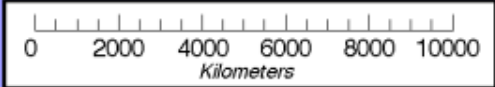
# Seeing a Solar Eclipse

- In order to witness a total solar eclipse you must be somewhere within the Moon's shadow at the time of New Moon.
- Because the Moon's shadow is relatively small, an eclipse is visible only in certain locations on Earth.
- The Moon's shadow moves rapidly over the surface of the Earth (up to 2000 mph).
- The maximum time possible for *totality* (when the Sun is totally eclipsed) is only about 6 minutes. But the entire process can take a couple of hours to observe.

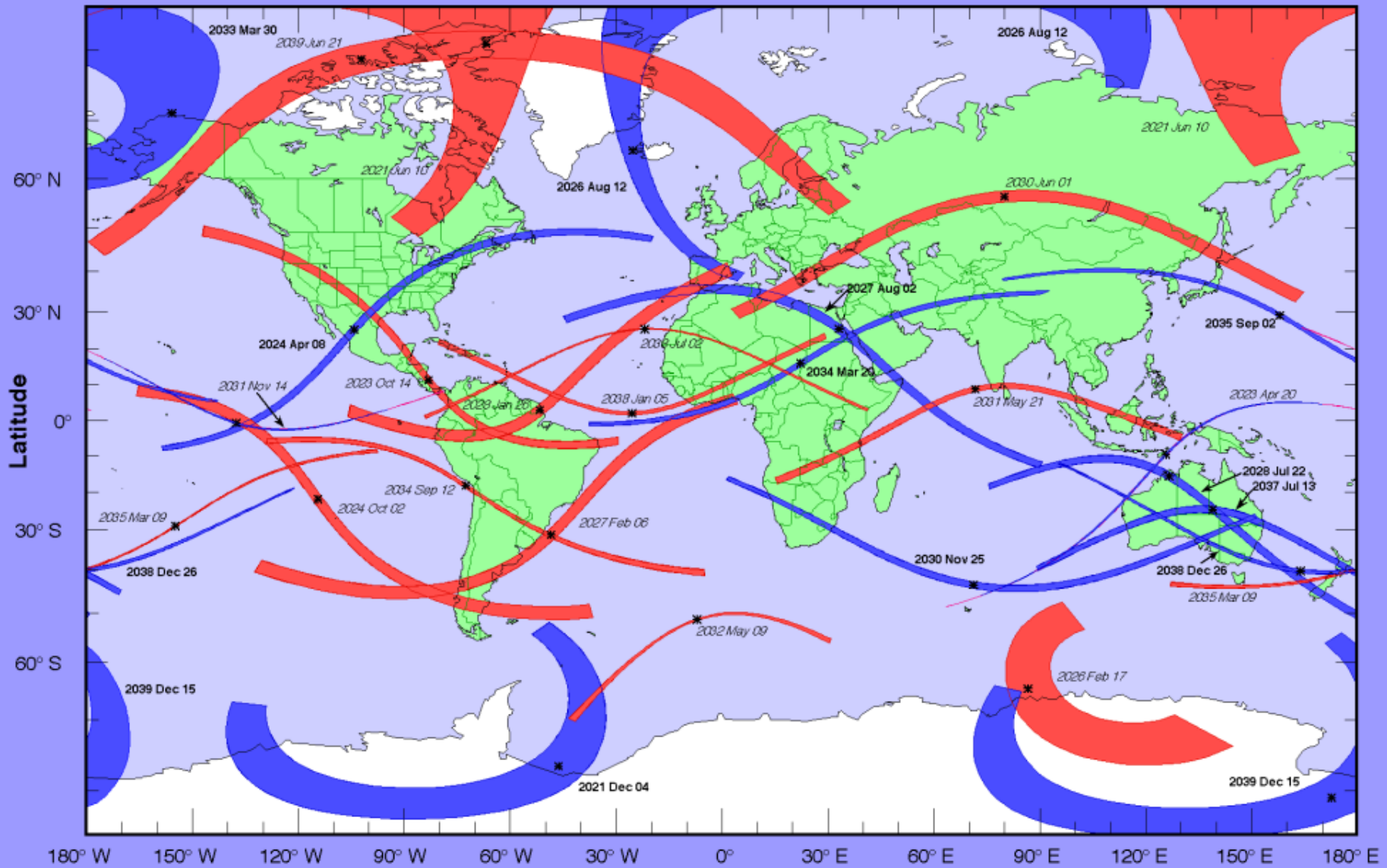
# Total and Annular Solar Eclipse Paths: 2001 – 2020



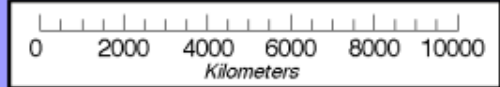
- Total Eclipse
- Annular Eclipse
- Hybrid Eclipse



# Total and Annular Solar Eclipse Paths: 2021 – 2040



<span style="color: blue;">■</span>	<b>Total Eclipse</b>
<span style="color: red;">■</span>	<b>Annular Eclipse</b>
<span style="color: magenta;">■</span>	<b>Hybrid Eclipse</b>



Partial

Total





[www.MrEclipse.com](http://www.MrEclipse.com)

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What phase is the Moon? Full!

Full Moon at *every* lunar eclipse? Yes!

Lunar eclipse at *every* Full Moon? No!



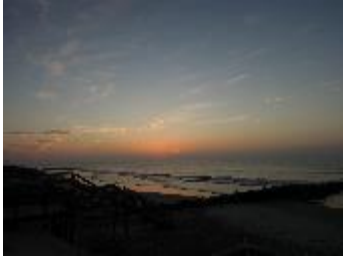
Penumbra Lunar Eclipse:  
Moon is slightly dimmed  
but misses the darkest  
part of Earth's shadow

Why the color (seen *only* during total)?

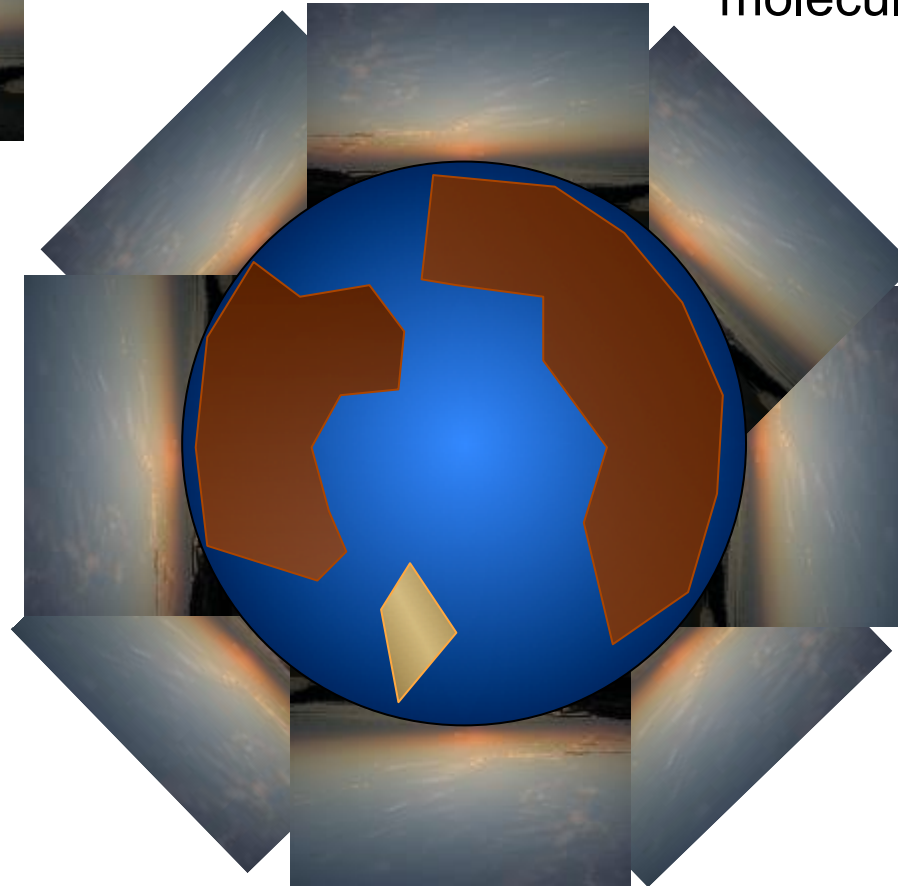




The orange/copper color of a total lunar eclipse is due to the same phenomenon as that which causes a sunset or sunrise to appear orange – the scattering of blue light by molecules in the atmosphere.



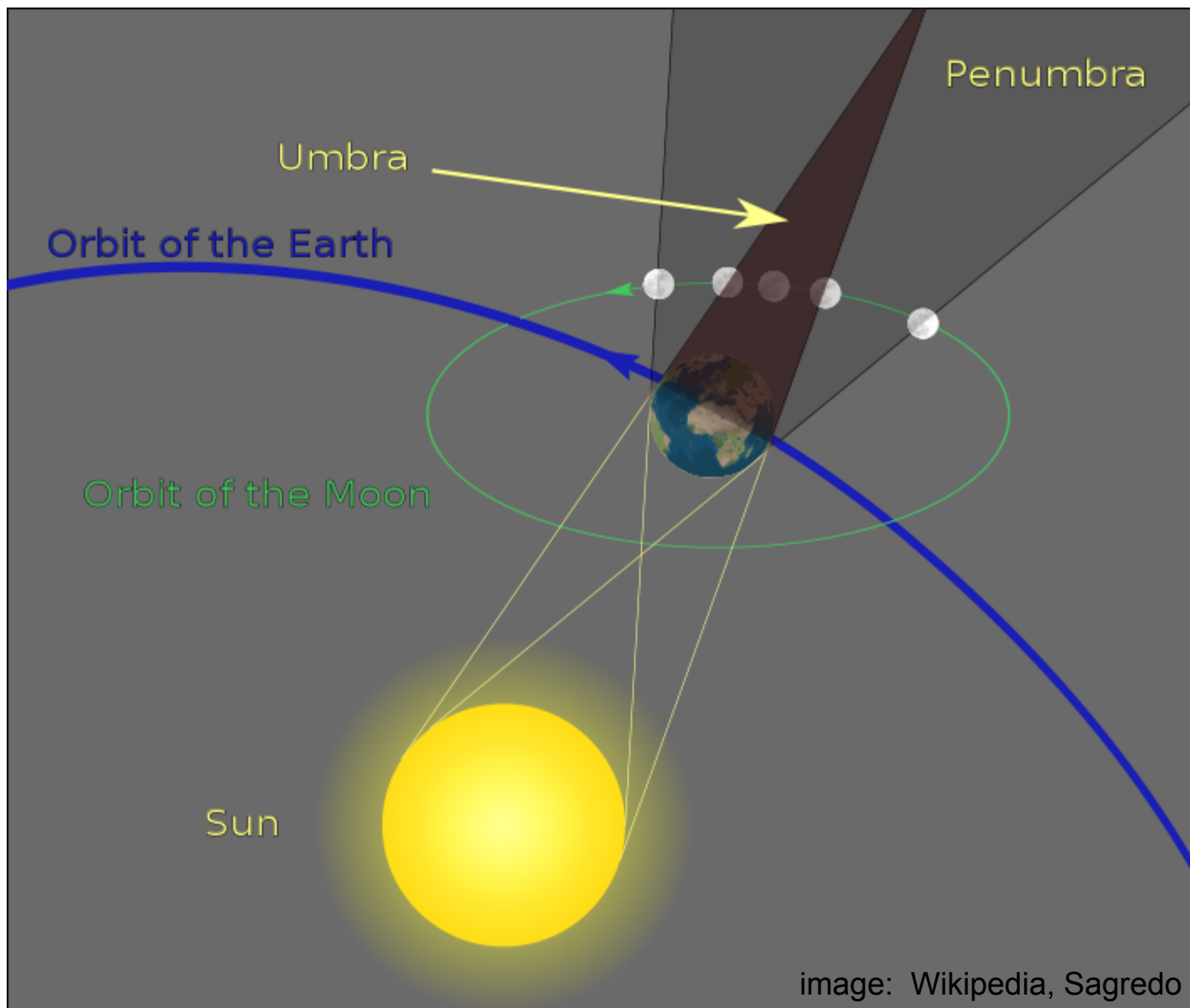
On one side of the Earth the Sun is setting...



...on the other the Sun is rising.

An observer on the Moon during a total lunar eclipse would see the Earth blocking the Sun, facing the night side of it. Light from the Sun bends through the atmosphere around all sides and illuminates the Moon.



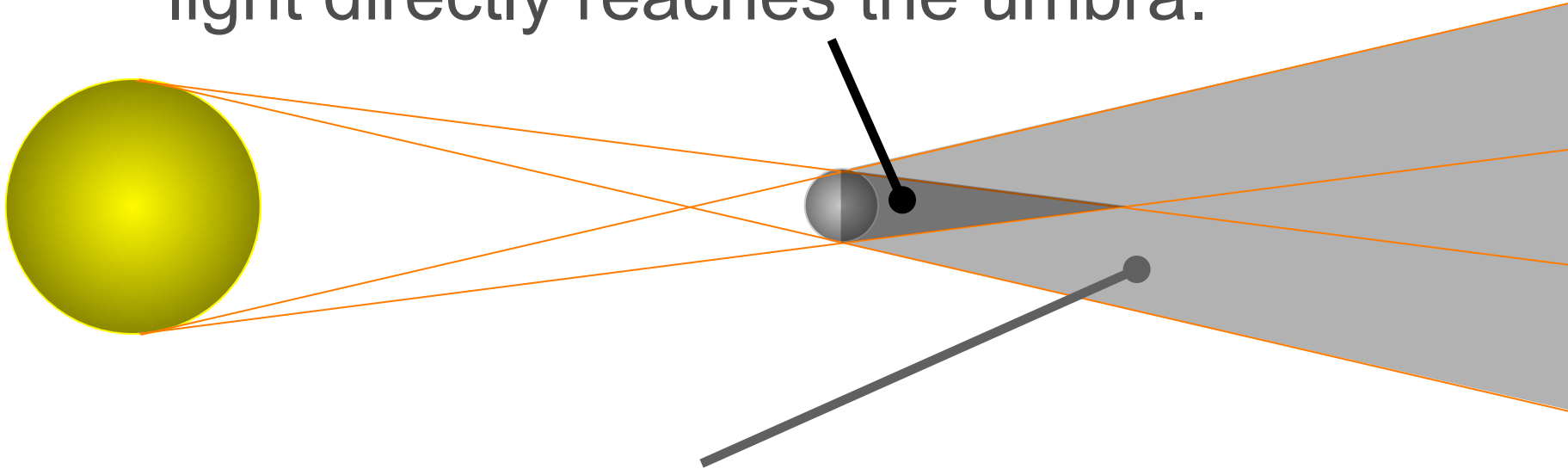


# Seeing a Lunar Eclipse

- In order to witness a lunar eclipse, you must be somewhere on the night side of Earth while the Full Moon is in the Earth's shadow.
- Because the Earth's shadow is relatively large a lunar eclipse progresses slowly.
- A lunar eclipse can last several hours from beginning to end.
- A penumbral eclipse causes such a slight change in the Moon's brightness that it is nearly impossible to notice.

# The Science of Shadows!

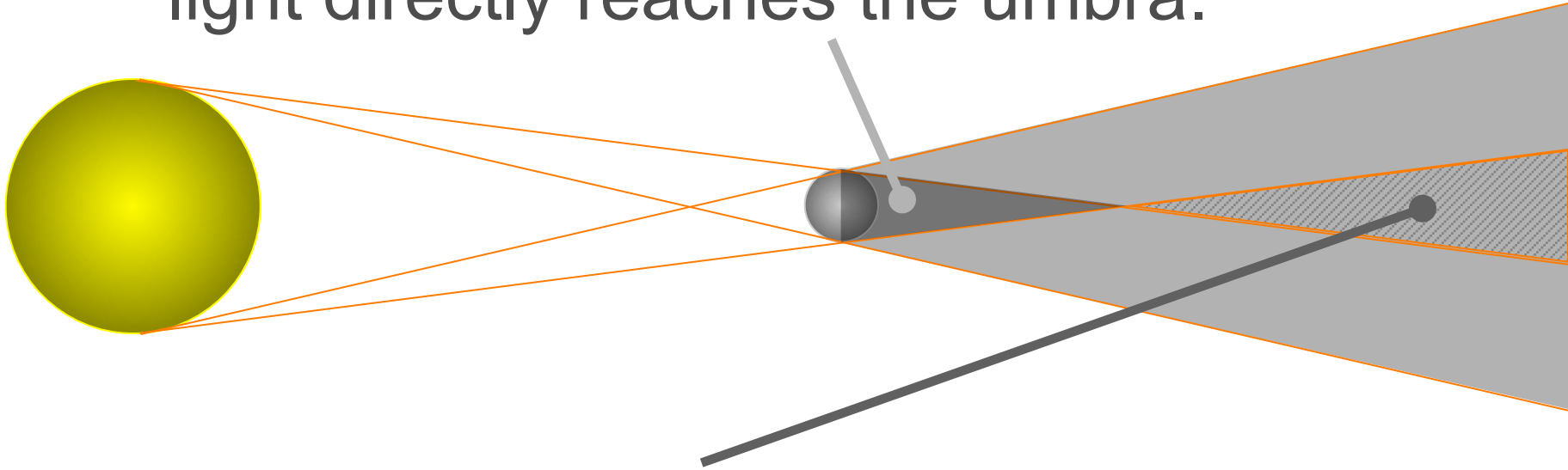
The **umbra** is the darkest part of the shadow. None of the Sun's light directly reaches the umbra.



The **penumbra** is a “partial shadow”. Some, but not all, of the Sun's light reaches the penumbra.

# The Science of Shadows!

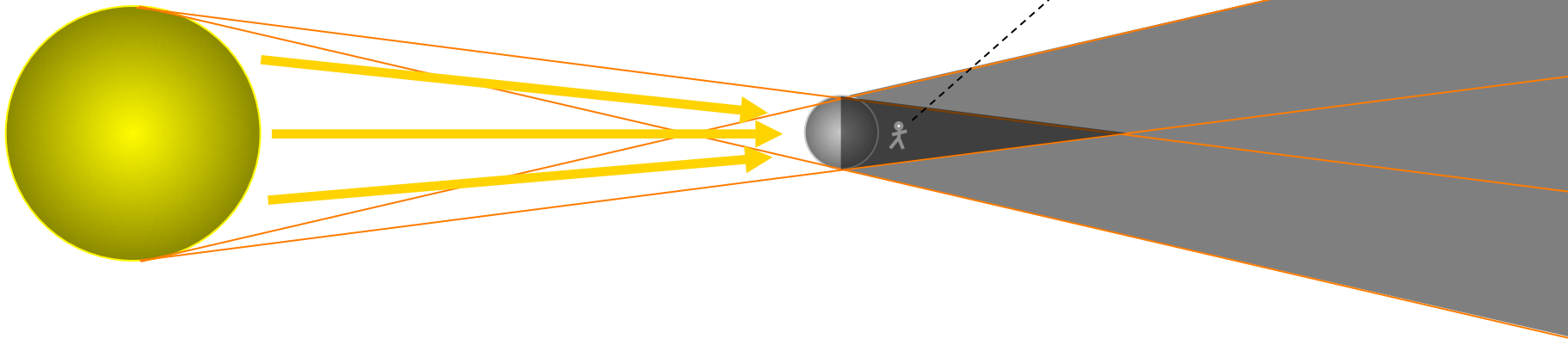
The **umbra** is the darkest part of the shadow. None of the Sun's light directly reaches the umbra.



The **antumbra** is the “inverse” of the umbra's cone. Like the penumbra, it is also a “partial shadow”.

# The Science of Shadows!

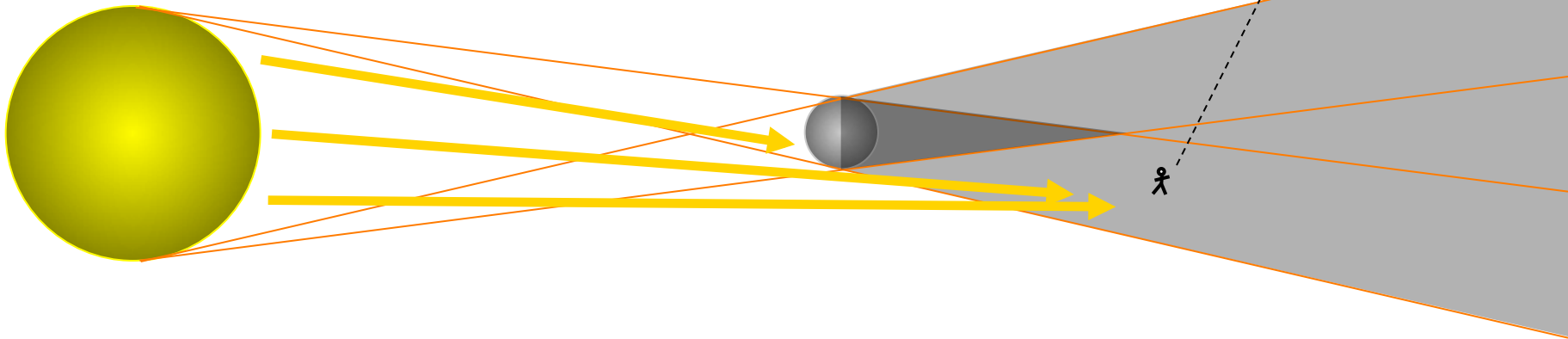
*All* parts of the Sun are blocked from view in the umbra.



# The Science of Shadows!



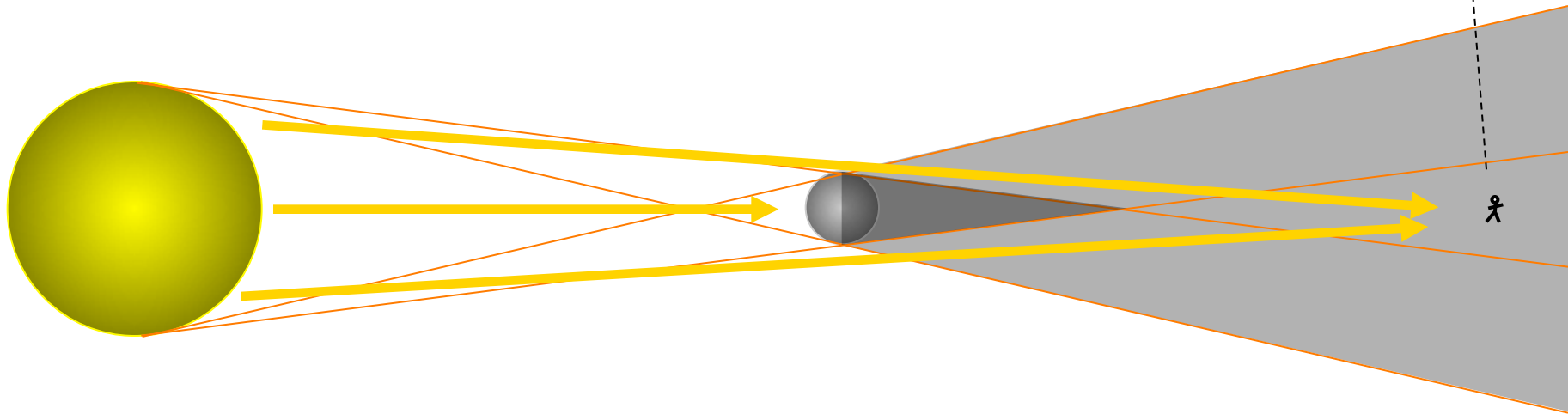
*All* parts of the Sun are blocked from view in the umbra.



Only *part* of the Sun is blocked from view in the penumbra.

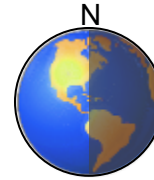
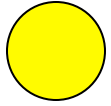
# The Science of Shadows!

Only the *middle* part of the Sun is blocked from view in the antumbra.

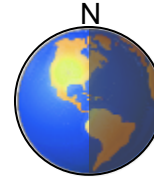
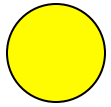


Only *part* of the Sun is blocked from view in the penumbra.

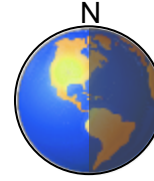
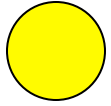
1.



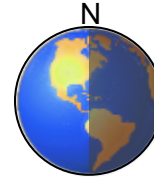
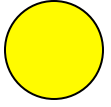
2.



3.



4.



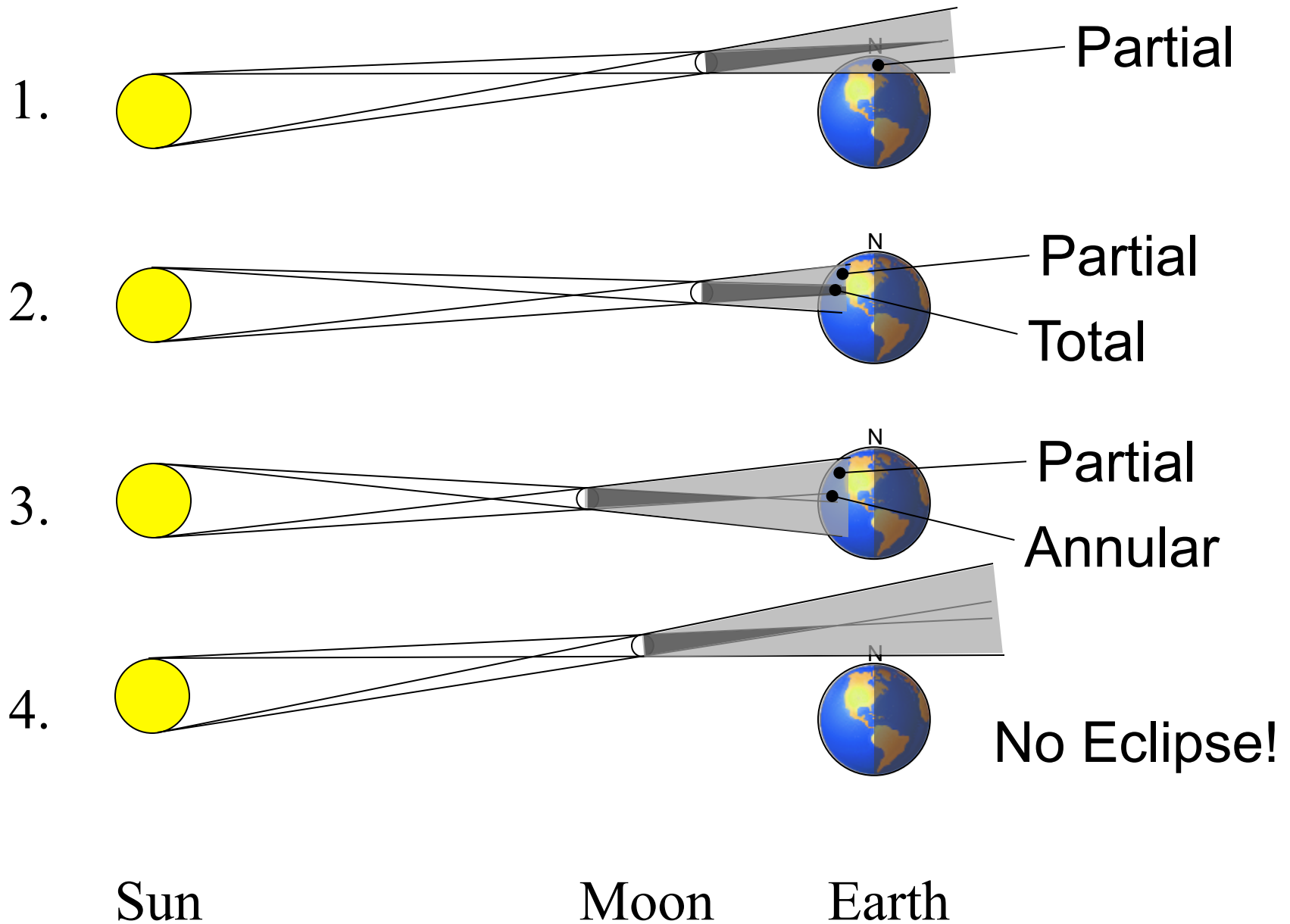
Sun

Moon

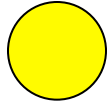
Earth



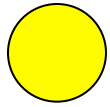
# Solar Eclipses



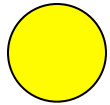
5.



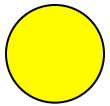
6.



7.



8.



Sun

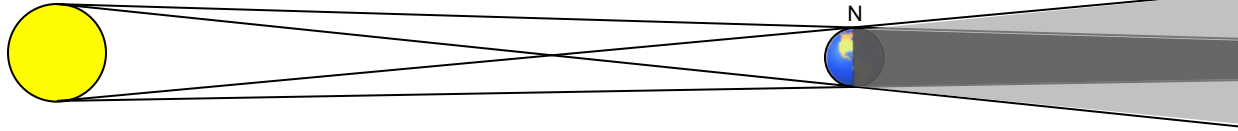
Earth

Moon

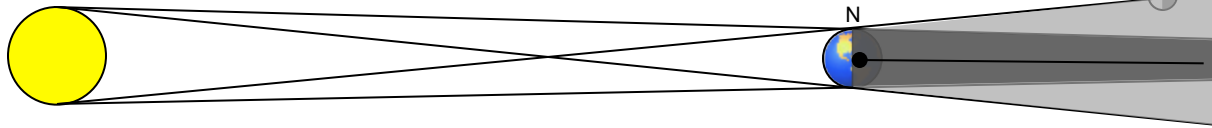
# Lunar Eclipses

No Eclipse!

5.

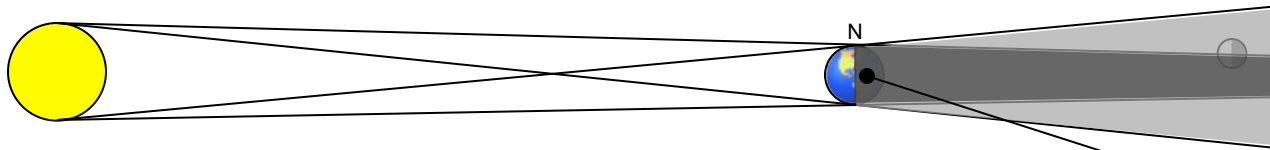


6.



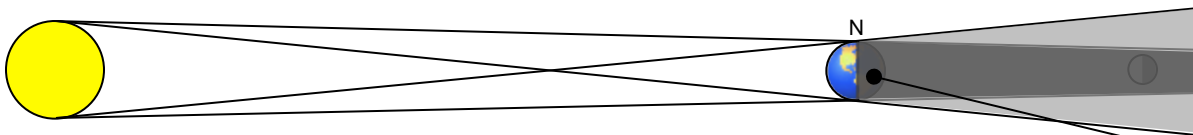
Penumbral

7.



Partial

8.

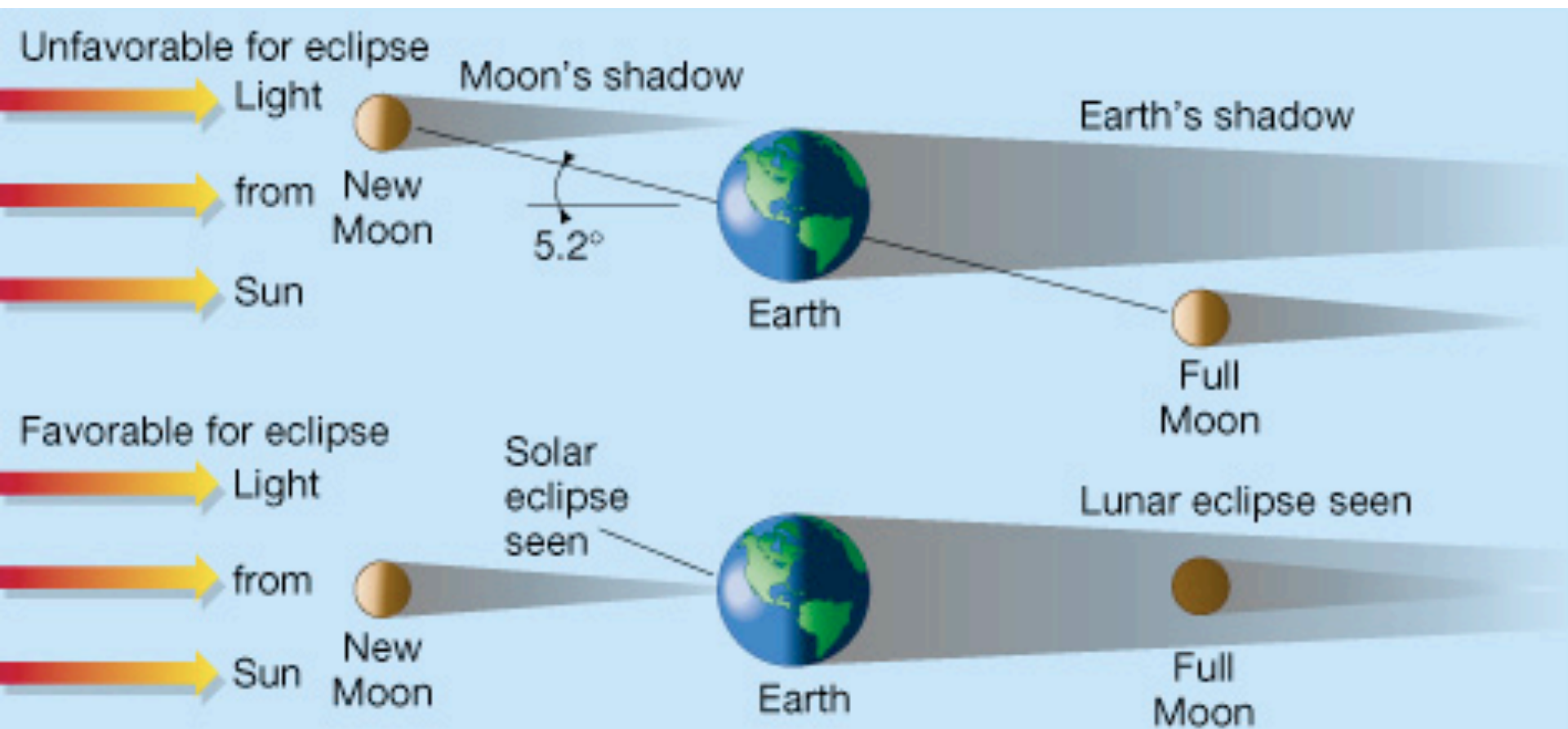


Total

Sun

Earth

Moon



# Frequency of Eclipses

- Approximately how often does a solar eclipse occur? How about a lunar eclipse?
- Consider sets of eclipses that occur within the same season of the same year – what patterns do you notice?
- Why do these patterns exist? Explain and relate to: ecliptic, synodic month, tropical year, Moon phases & age, etc.

2007	Mar 3	Lun – Tot
	<b>Mar 19</b>	<b>Sol – Part</b>
	Aug 28	Lun – Tot
	<b>Sep 11</b>	<b>Sol – Part</b>
2008	<b>Feb 7</b>	<b>Sol – Ann</b>
	Feb 21	Lun – Tot
	<b>Aug 1</b>	<b>Sol – Tot</b>
	Aug 16	Lun – Part
2009	<b>Jan 26</b>	<b>Sol – Ann</b>
	Feb 9	Lun – Pen
	July 7	Lun – Pen
	<b>Jul 22</b>	<b>Sol Total</b>
	Aug 6	Lun – Pen

	Dec 31	Lun – Part
2010	<b>Jan 15</b>	<b>Sol – Ann</b>
	Jun 26	Lun – Part
	<b>Jul 11</b>	<b>Sol – Tot</b>
	Dec 21	Lun – Tot
2011	<b>Jan 4</b>	<b>Sol – Part</b>
	<b>Jun 1</b>	<b>Sol – Part</b>
	Jun 15	Lun – Tot
	<b>Jul 1</b>	<b>Sol – Part</b>
	<b>Nov 25</b>	<b>Sol – Part</b>
	Dec 10	Lun – Tot
2012	<b>May 20</b>	<b>Sol – Ann</b>
	Jun 4	Lun – Part

	<b>Nov 13</b>	<b>Sol – Tot</b>
	Nov 28	Lun – Pen
2013	Apr 25	Lun – Part
	<b>May 10</b>	<b>Sol – Ann</b>
	May 25	Lun – Pen
	Oct 18	Lun – Pen
	<b>Nov 3</b>	<b>Sol – Hyb</b>
2014	Apr 15	Lun – Tot
	<b>Apr 29</b>	<b>Sol – Ann</b>
	Oct 8	Lun – Tot
	<b>Oct 23</b>	<b>Sol – Part</b>
2015	<b>Mar 20</b>	<b>Sol – Tot</b>
	Apr 4	Lun – Part

	<b>Sep 13</b>	<b>Sol – Part</b>
	Sep 28	Lun – Tot
2016	<b>Mar 9</b>	<b>Sol – Tot</b>
	Mar 23	Lun – Pen
	<b>Sep 1</b>	<b>Sol – Ann</b>
	Sep 16	Lun – Pen
2017	Feb 11	Lun – Pen
	<b>Feb 26</b>	<b>Sol – Ann</b>
	Aug 7	Lun – Part
	<b>Aug 21</b>	<b>Sol – Tot</b>
2018	Jan 31	Lun – Tot
	<b>Feb 15</b>	<b>Sol – Part</b>
	<b>Jul 13</b>	<b>Sol – Part</b>

2021	May 26	Lun – Tot
	<b>Jun 10</b>	<b>Sol – Ann</b>
	Nov 19	Lun – Part
	<b>Dec 4</b>	<b>Sol – Tot</b>
2022	<b>Apr 30</b>	<b>Sol – Part</b>
	May 16	Lun – Tot
	<b>Oct 25</b>	<b>Sol – Part</b>
	Nov 8	Lun – Tot
2023	<b>Apr 20</b>	<b>Sol – Hyb</b>
	May 5	Lun – Pen
	<b>Oct 14</b>	<b>Sol – Ann</b>
	Oct 28	Lun - Part

2024	Mar 25	Lun – Pen
	<b>Apr 8</b>	<b>Sol – Tot</b>
	Sep 18	Lun – Part
	<b>Oct 2</b>	<b>Sol – Ann</b>
2025	Mar 14	
	<b>Mar 29</b>	
	Sep 7	
	<b>Sep 21</b>	
2026	<b>Feb 17</b>	<b>Sol – Ann</b>
		Lun – Tot
		<b>Sol – Tot</b>
		Lun – Part



2021	May 26	Lun – Tot
	<b><i>Jun 10</i></b>	<b><i>Sol – Ann</i></b>
	Nov 19	Lun – Part
	<b><i>Dec 4</i></b>	<b><i>Sol – Tot</i></b>
2022	<b><i>Apr 30</i></b>	<b><i>Sol – Part</i></b>
	May 16	Lun – Tot
	<b><i>Oct 25</i></b>	<b><i>Sol – Part</i></b>
	Nov 8	Lun – Tot
2023	<b><i>Apr 20</i></b>	<b><i>Sol – Hyb</i></b>
	May 5	Lun – Pen
	<b><i>Oct 14</i></b>	<b><i>Sol – Ann</i></b>
	Oct 28	Lun - Part

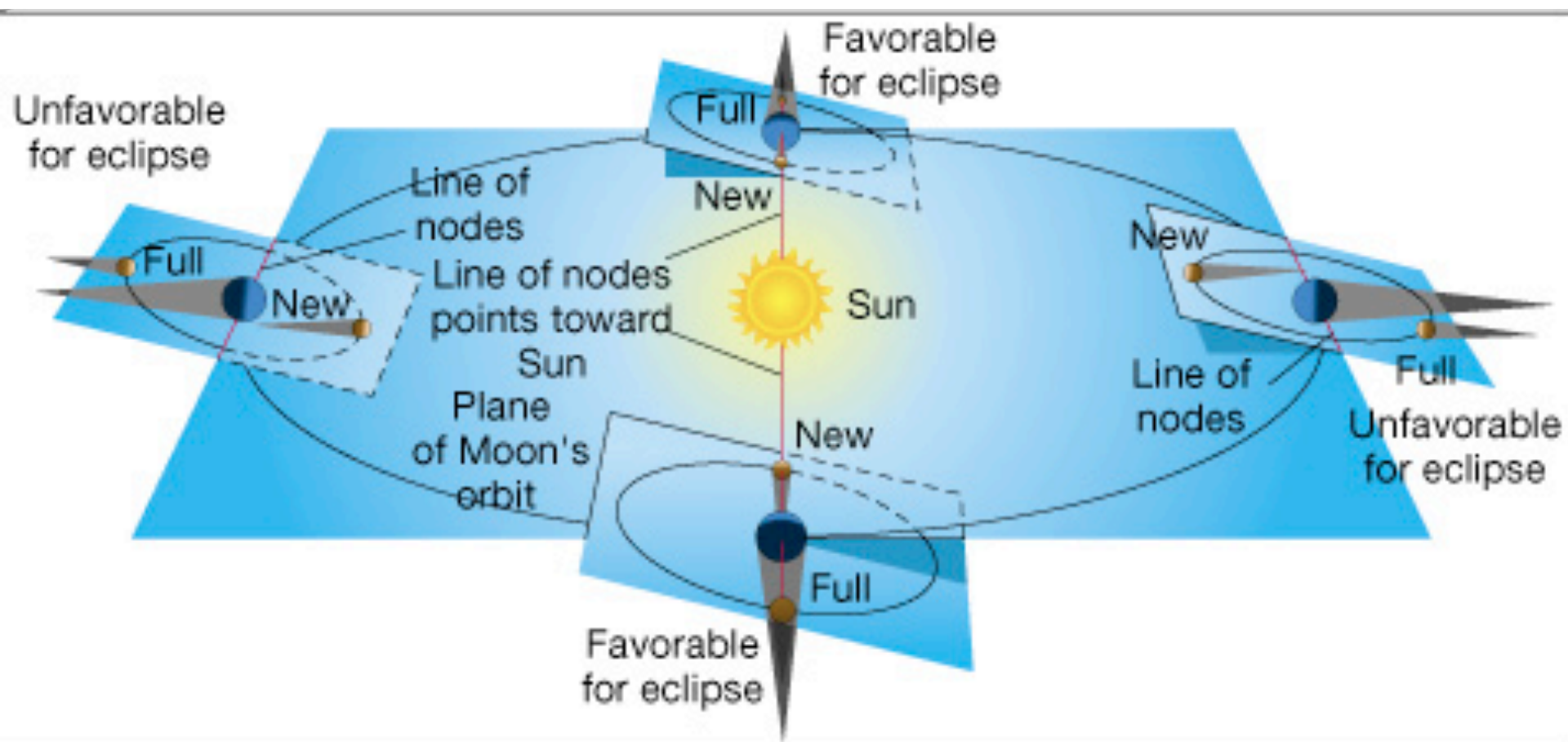
2024	Mar 25	Lun – Pen
	<b><i>Apr 8</i></b>	<b><i>Sol – Tot</i></b>
	Sep 18	Lun – Part
	<b><i>Oct 2</i></b>	<b><i>Sol – Ann</i></b>
2025	Mar 14	Lun – Tot
	<b><i>Mar 29</i></b>	<b><i>Sol – Part</i></b>
	Sep 7	Lun – Tot
	<b><i>Sep 21</i></b>	<b><i>Sol – Part</i></b>
2026	<b><i>Feb 17</i></b>	<b><i>Sol – Ann</i></b>
	Mar 3	Lun – Tot
	<b><i>Aug 12</i></b>	<b><i>Sol – Tot</i></b>
	Aug 28	Lun – Part

# Frequency of Solar Eclipses

- A solar eclipse can *only* occur during a *new* Moon, but does not occur at *every* new Moon.
- In order to occur, the Moon must be new at about the same time it is crossing the ecliptic.
- There are typically two solar eclipses per year, occurring about 6 months apart.

# Frequency of Lunar Eclipses

- A lunar eclipse can *only* occur during a *full* Moon, but does not occur at *every* full Moon.
- In order to occur, the Moon must be full at about the same time it is crossing the ecliptic.
- There are typically two lunar eclipses per year, occurring about 6 months apart.
- A lunar eclipse is followed two weeks later by a solar eclipse, or vice versa – most often a pair with one of each, but sometimes a triplet. These are the alternating “eclipse seasons” shown by the shading in the preceding tables.



# A simplified 3-D view of Moon orbiting Earth while Earth orbits the Sun

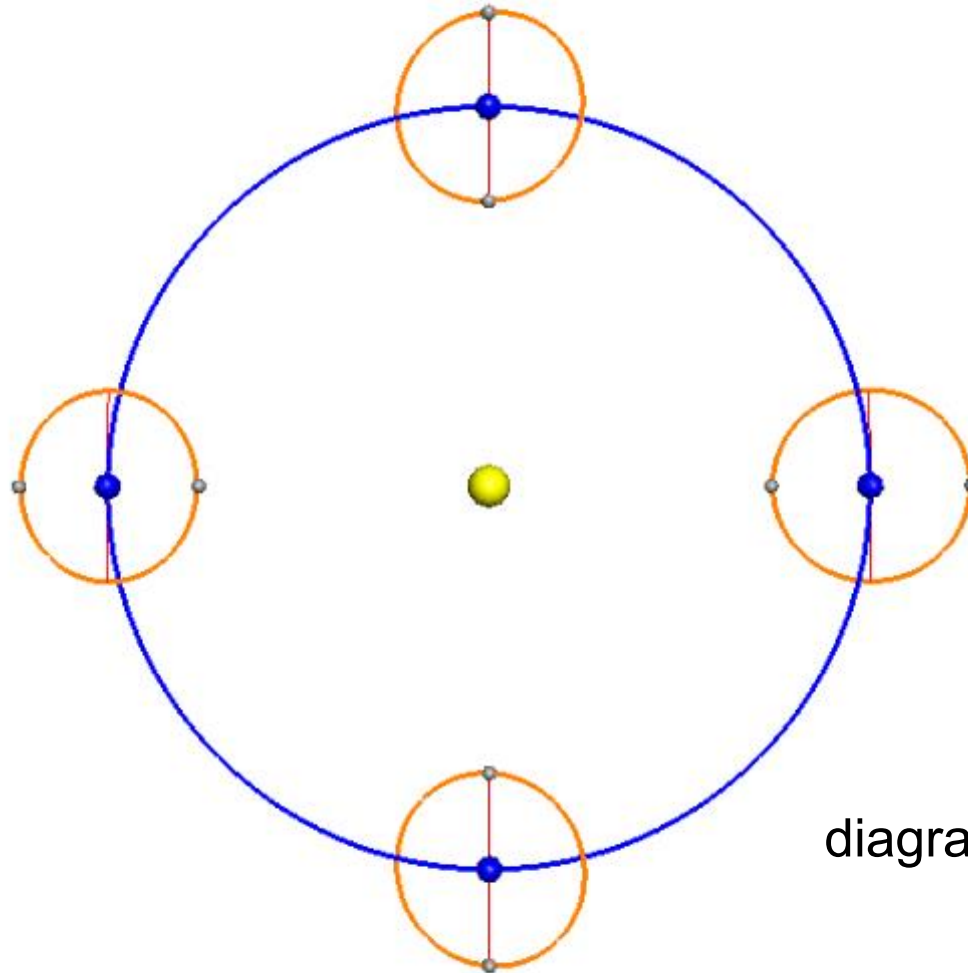
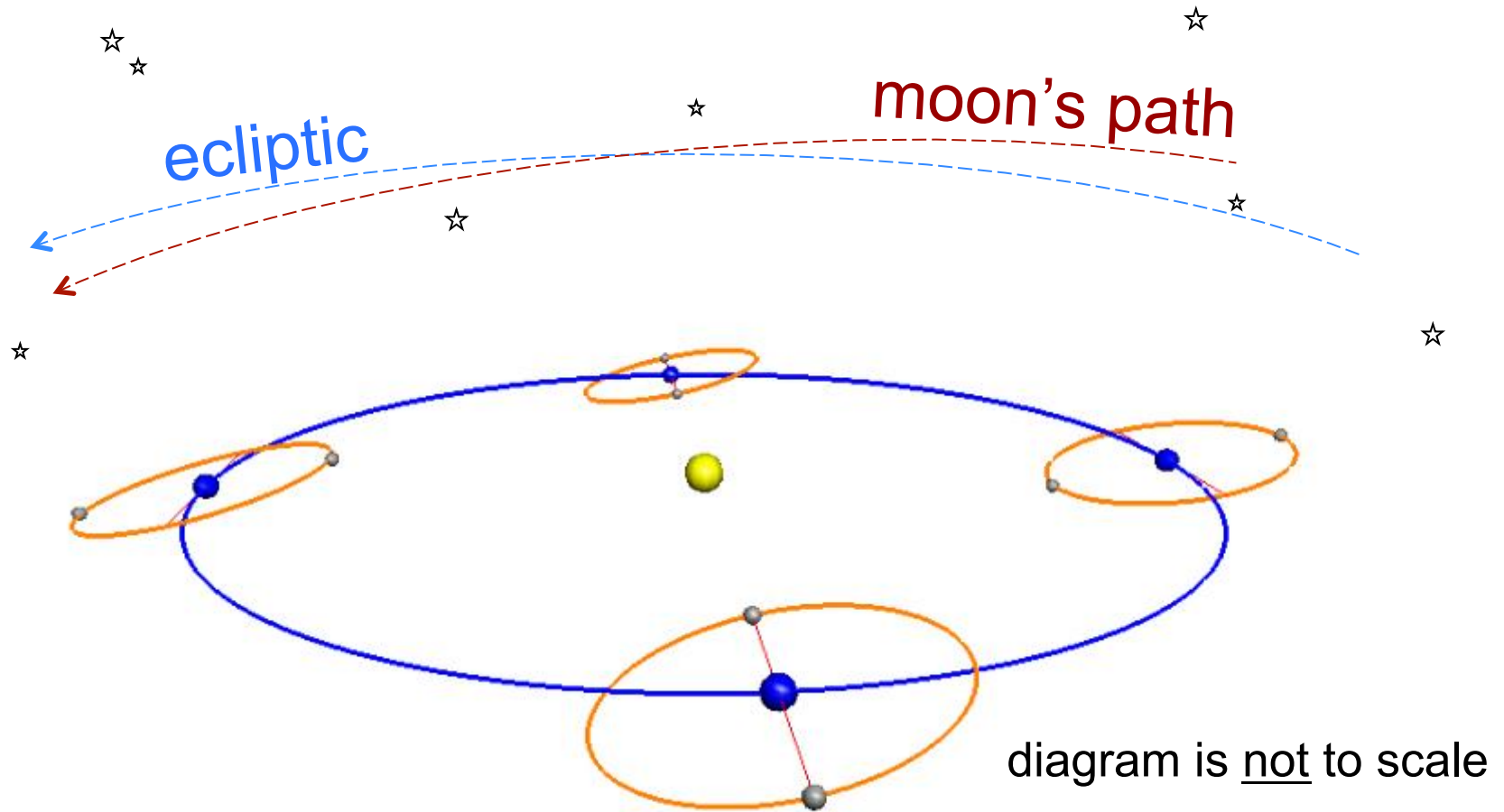


diagram is not to scale

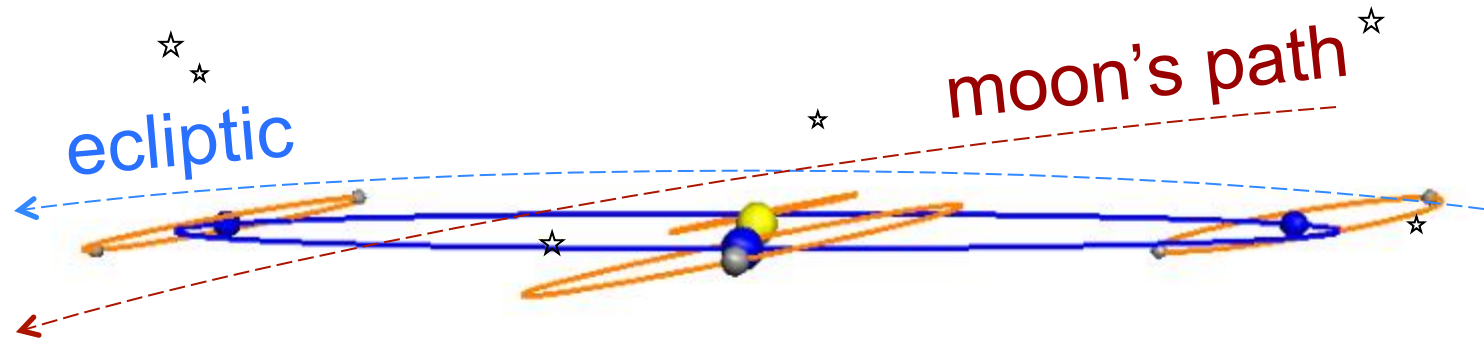
The red line shows the orientation of the intersection of the plane of Earth's orbit with the plane of Moon's orbit – this is called the “line of nodes”.

# A simplified 3-D view of Moon orbiting Earth while Earth orbits the Sun



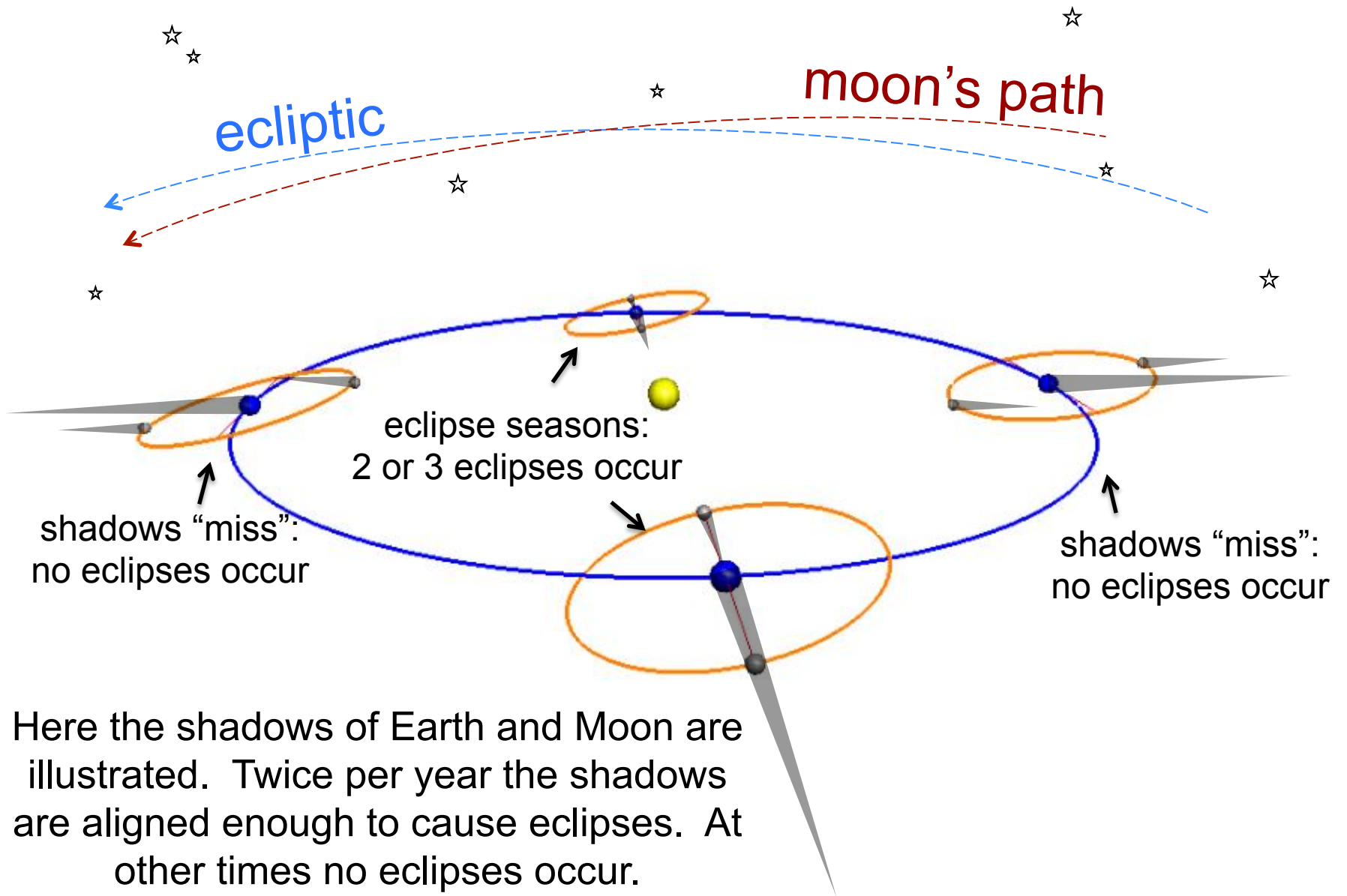
An observer on the Earth shown on “this side” of the orbit would see the ecliptic and Moon’s path on the celestial sphere “background of stars” as illustrated here.

# A simplified 3-D view of Moon orbiting Earth while Earth orbits the Sun



☆  
An observer on the Earth shown on “this side” of the orbit would see the ecliptic and Moon’s path on the celestial sphere “background of stars” as illustrated here.

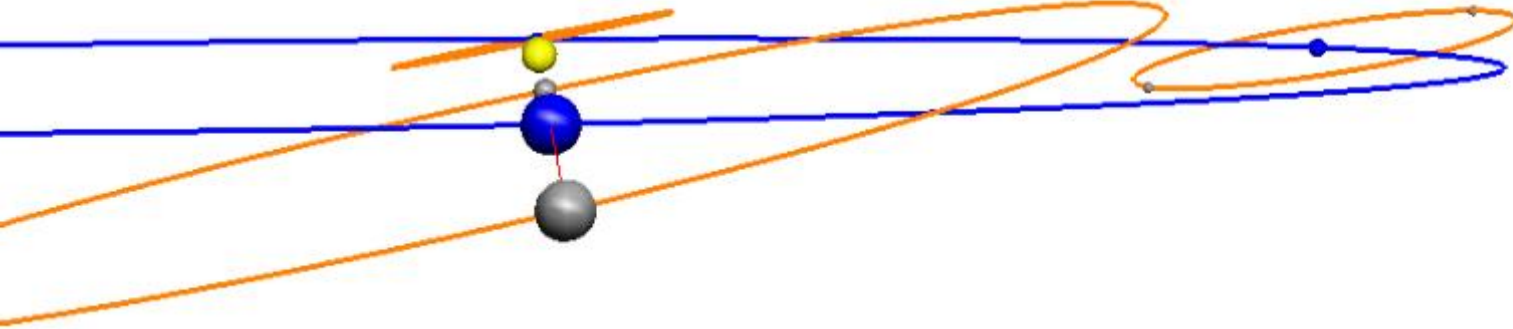
☆  
diagram is not to scale



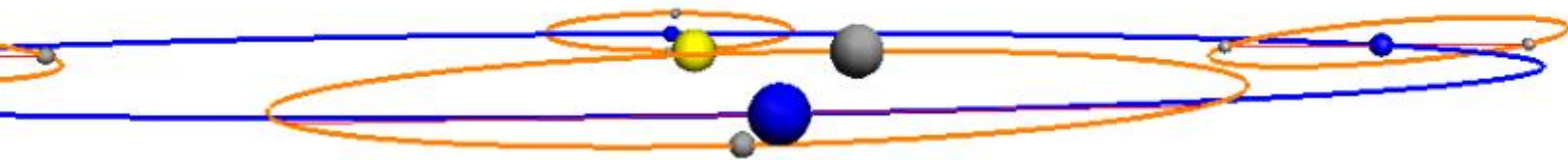
Here the shadows of Earth and Moon are illustrated. Twice per year the shadows are aligned enough to cause eclipses. At other times no eclipses occur.



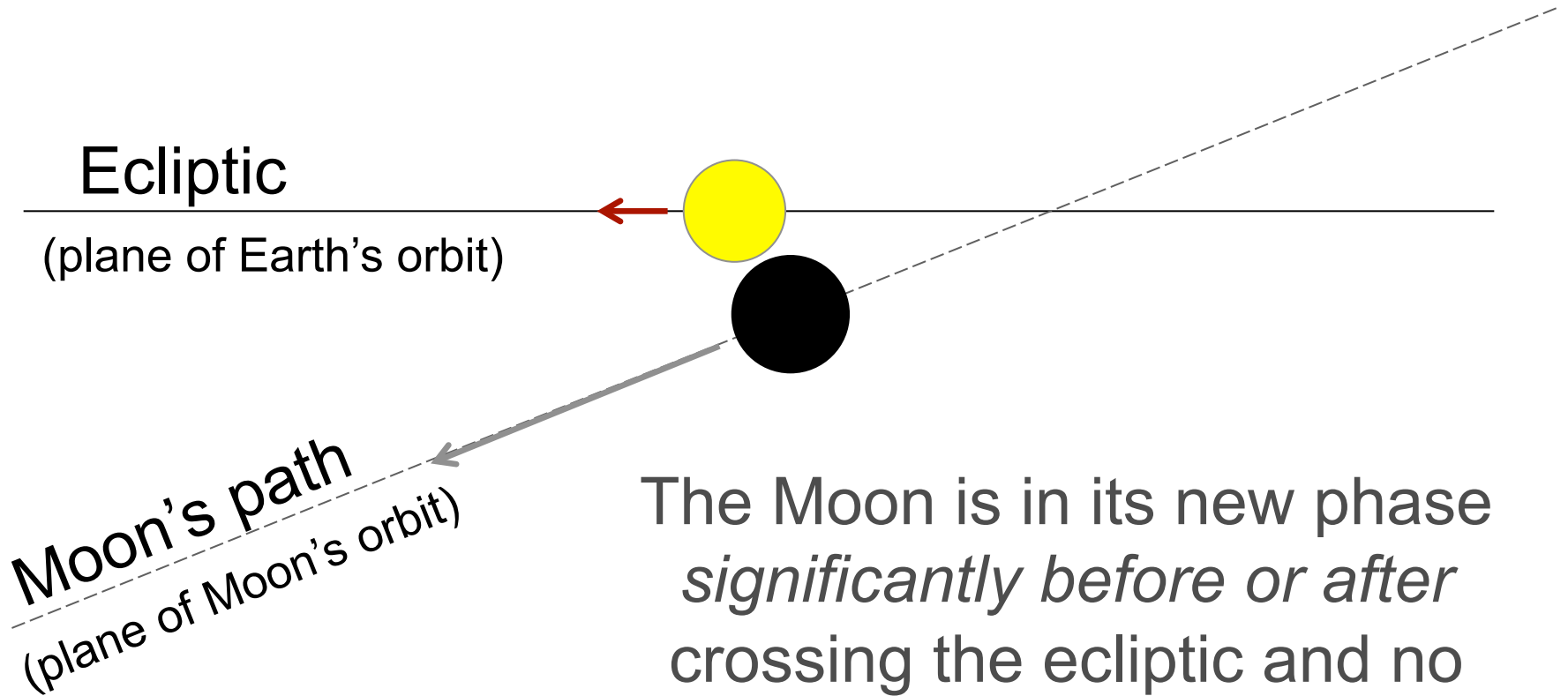
A simplified 3-D view of Moon orbiting Earth while Earth orbits the Sun



diagrams are not to scale

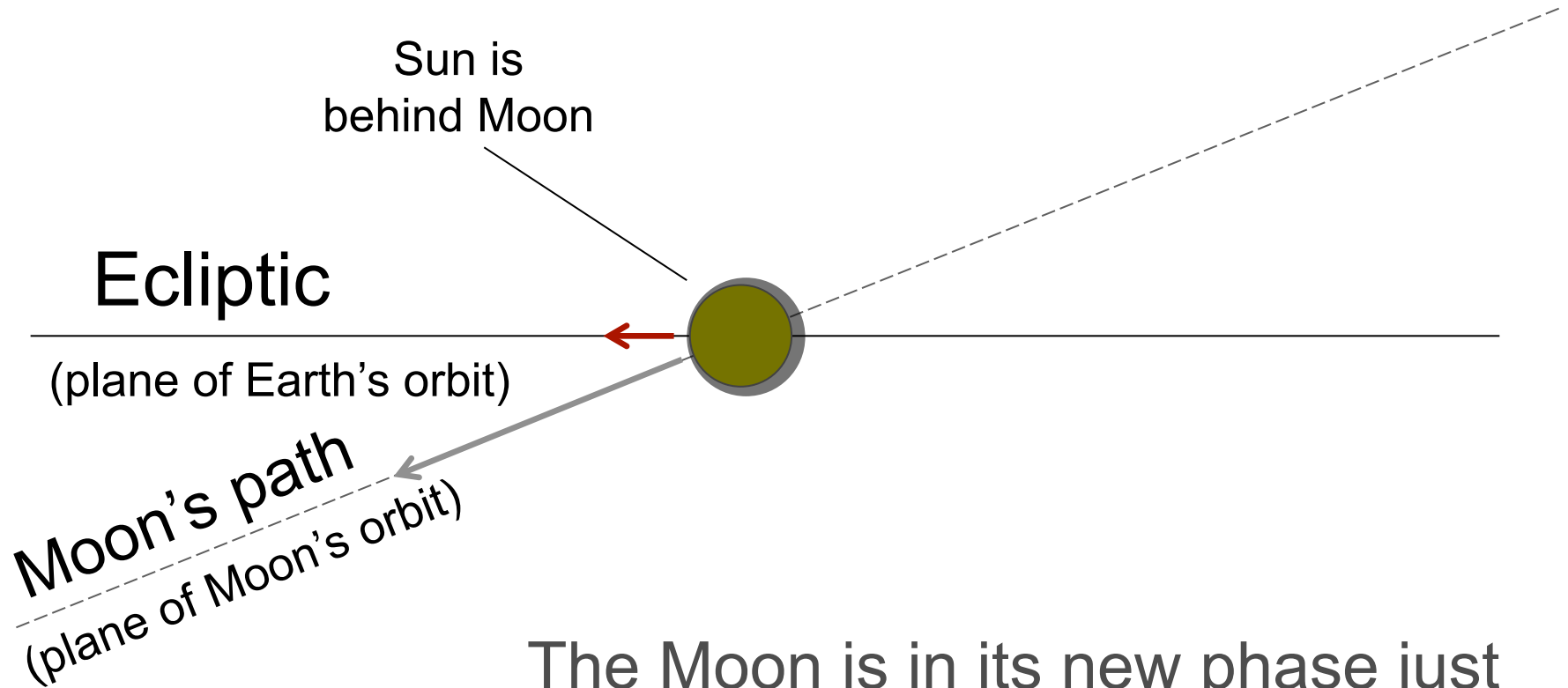


# No Eclipse!



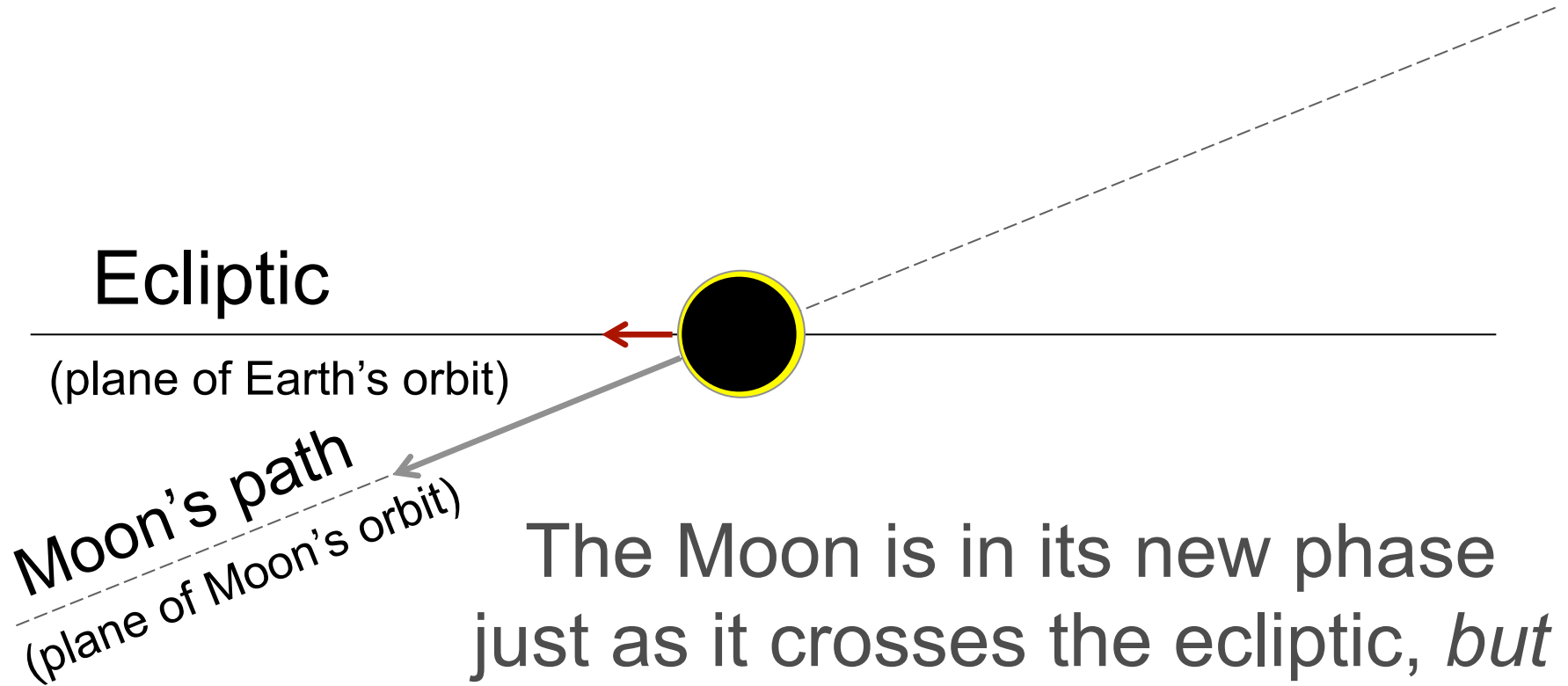
The Moon is in its new phase *significantly before or after* crossing the ecliptic and no part of the Sun is covered by it.

# A Total Solar Eclipse



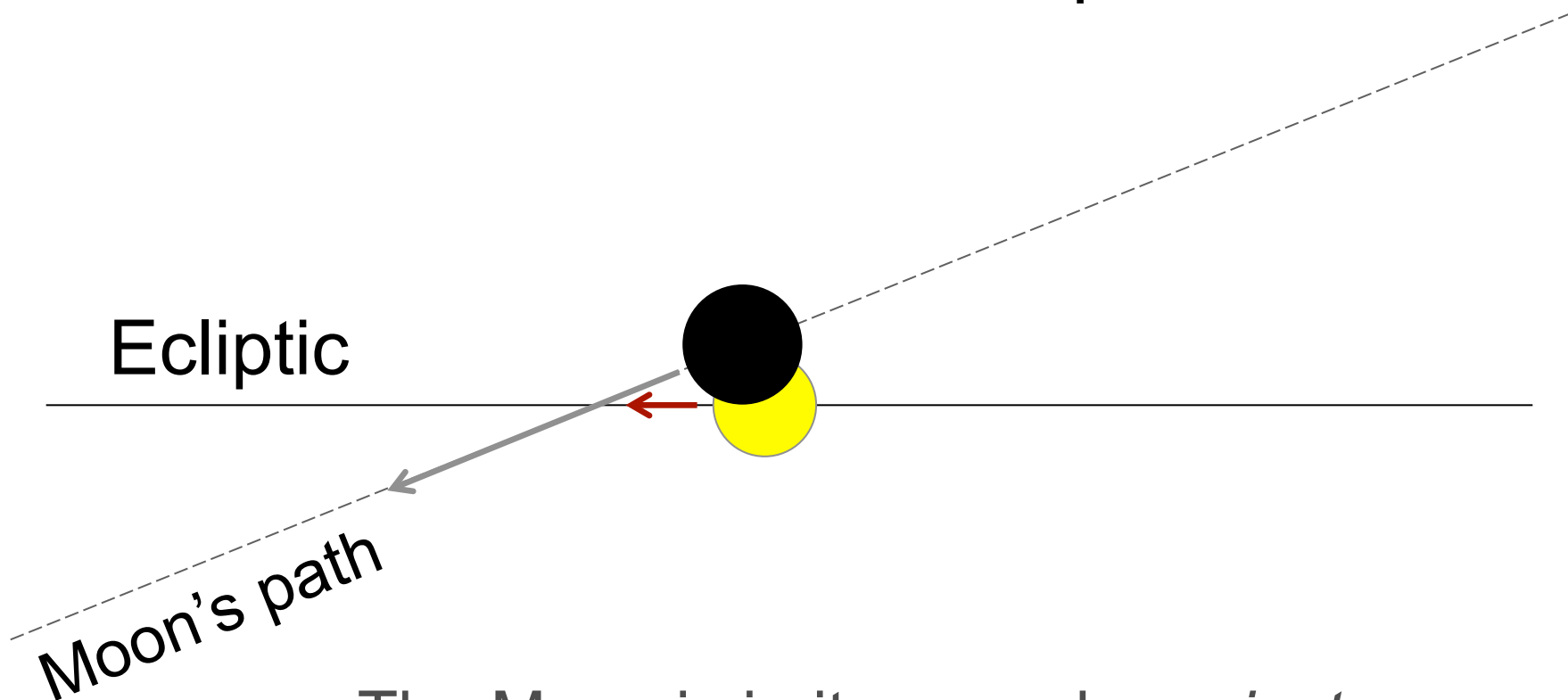
The Moon is in its new phase just as it crosses the ecliptic and at some point completely obscures the bright surface of the Sun.

# An Annular Solar Eclipse



The Moon is in its new phase just as it crosses the ecliptic, *but* the Sun's angular diameter is greater and at a certain point it appears like a bright ring.

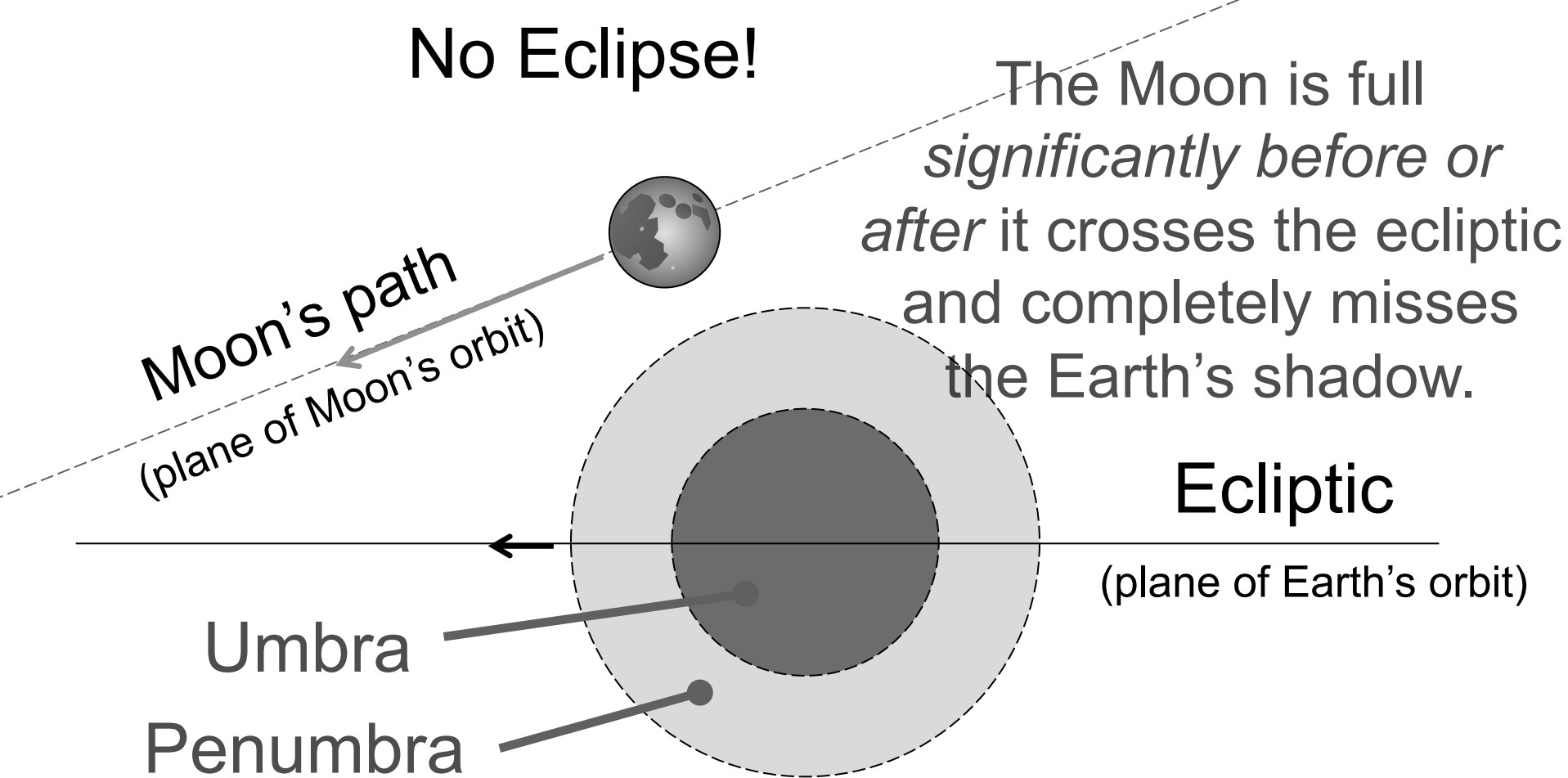
# A Partial Solar Eclipse



The Moon is in its new phase *just before or just after* crossing the ecliptic and obscures only part of the bright surface of the Sun.

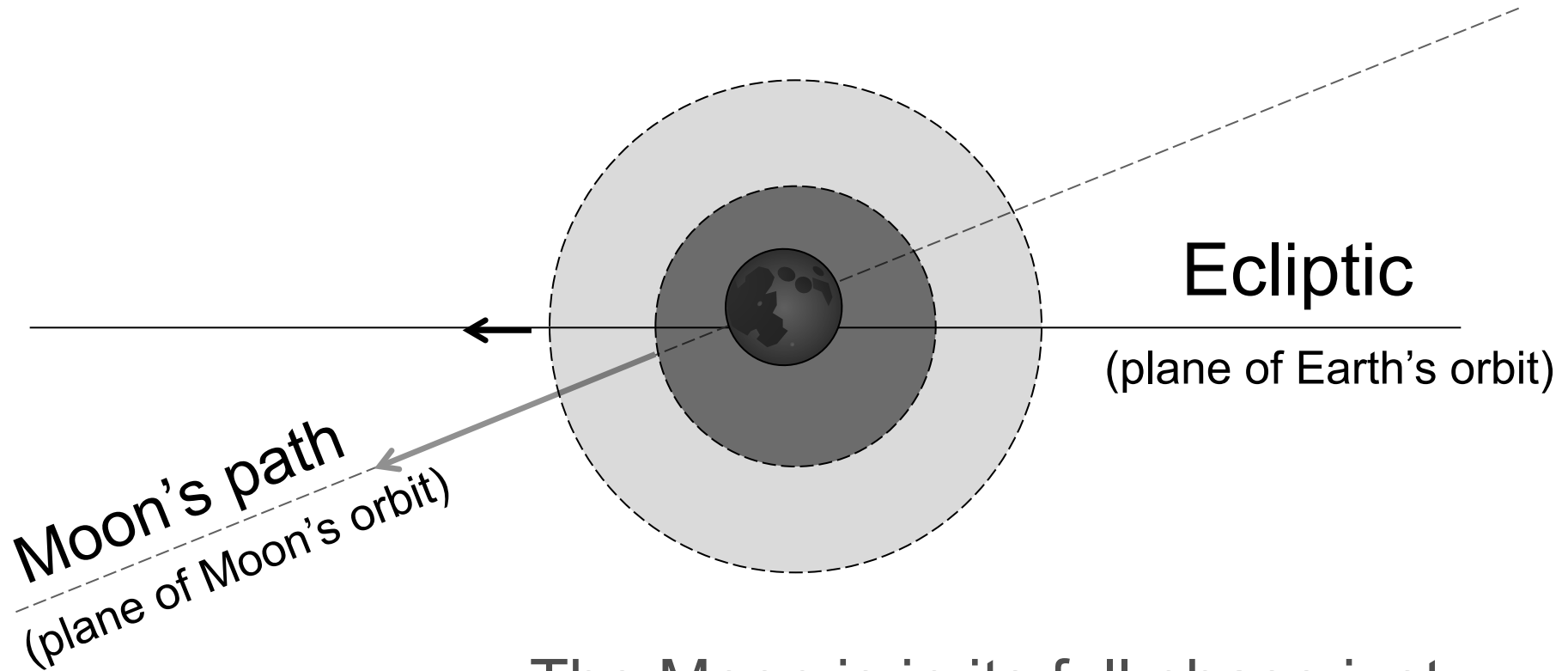
# No Eclipse!

The Moon is full significantly before or after it crosses the ecliptic and completely misses the Earth's shadow.



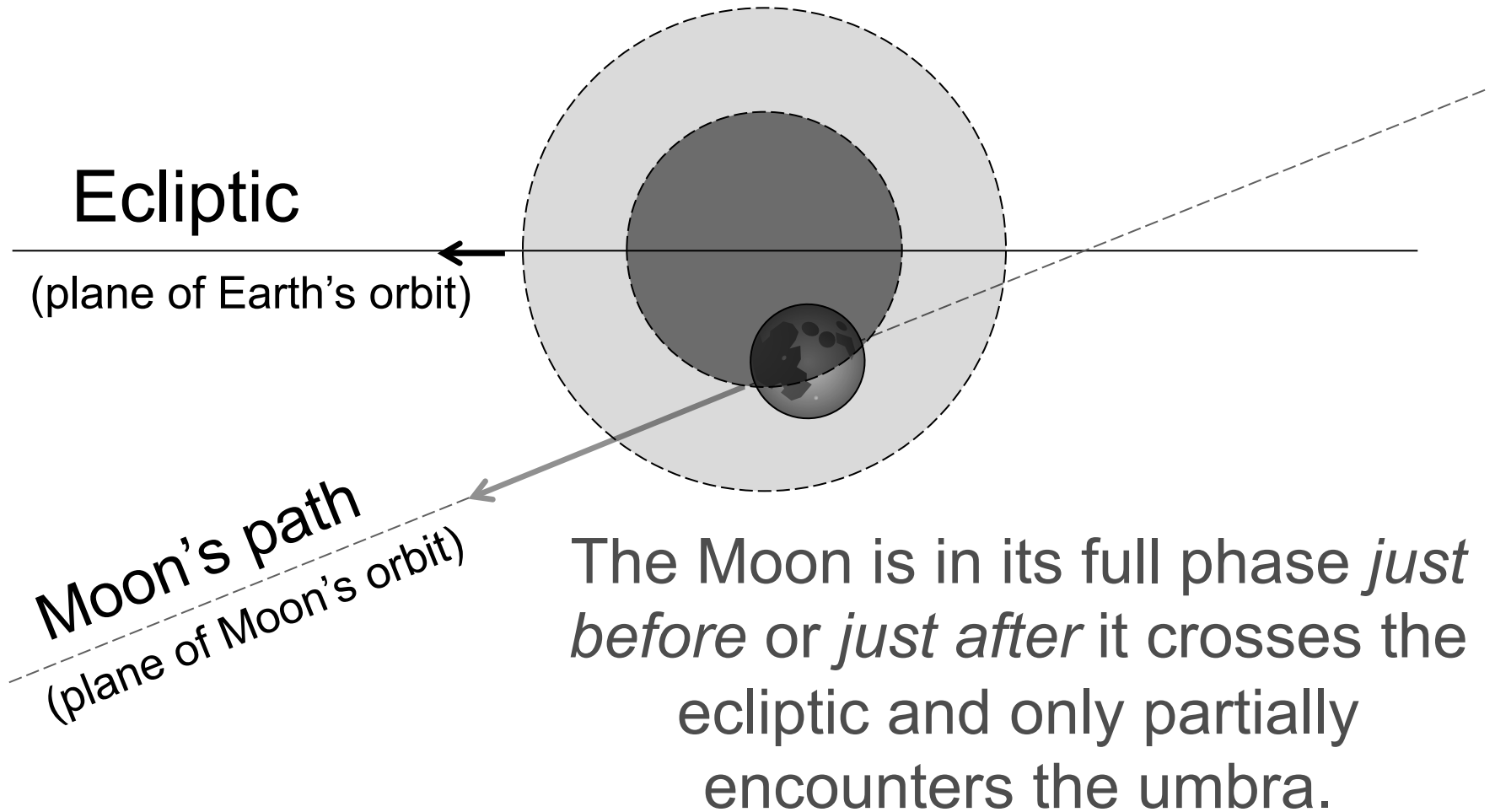
The Earth's shadow always lies in the plane of its orbit. Therefore its umbra and penumbra are always located centered on the ecliptic exactly opposite the Sun's location on the ecliptic. As the Sun traverses the ecliptic, so too does the Earth's shadow, in lock step on the opposite side of the ecliptic.

# A Total Lunar Eclipse



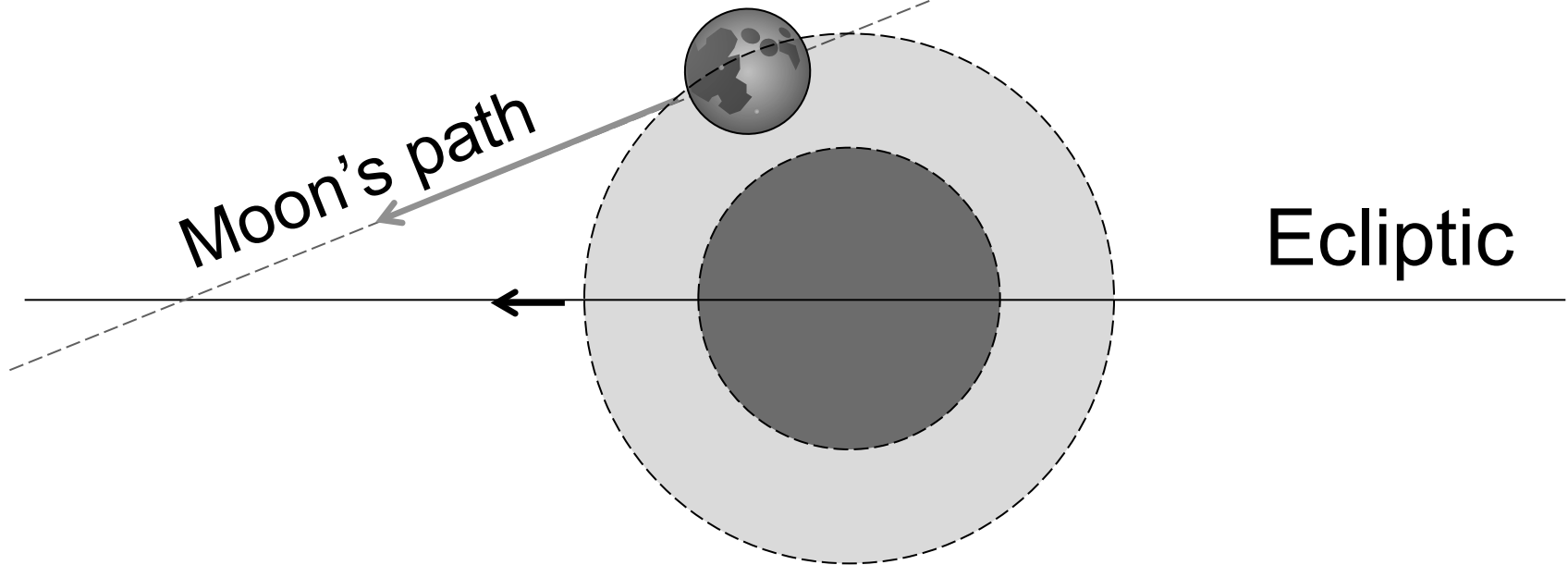
The Moon is in its full phase just as it crosses the ecliptic and is fully enveloped in the umbra.

# A Partial Lunar Eclipse





# A Penumbral Lunar Eclipse



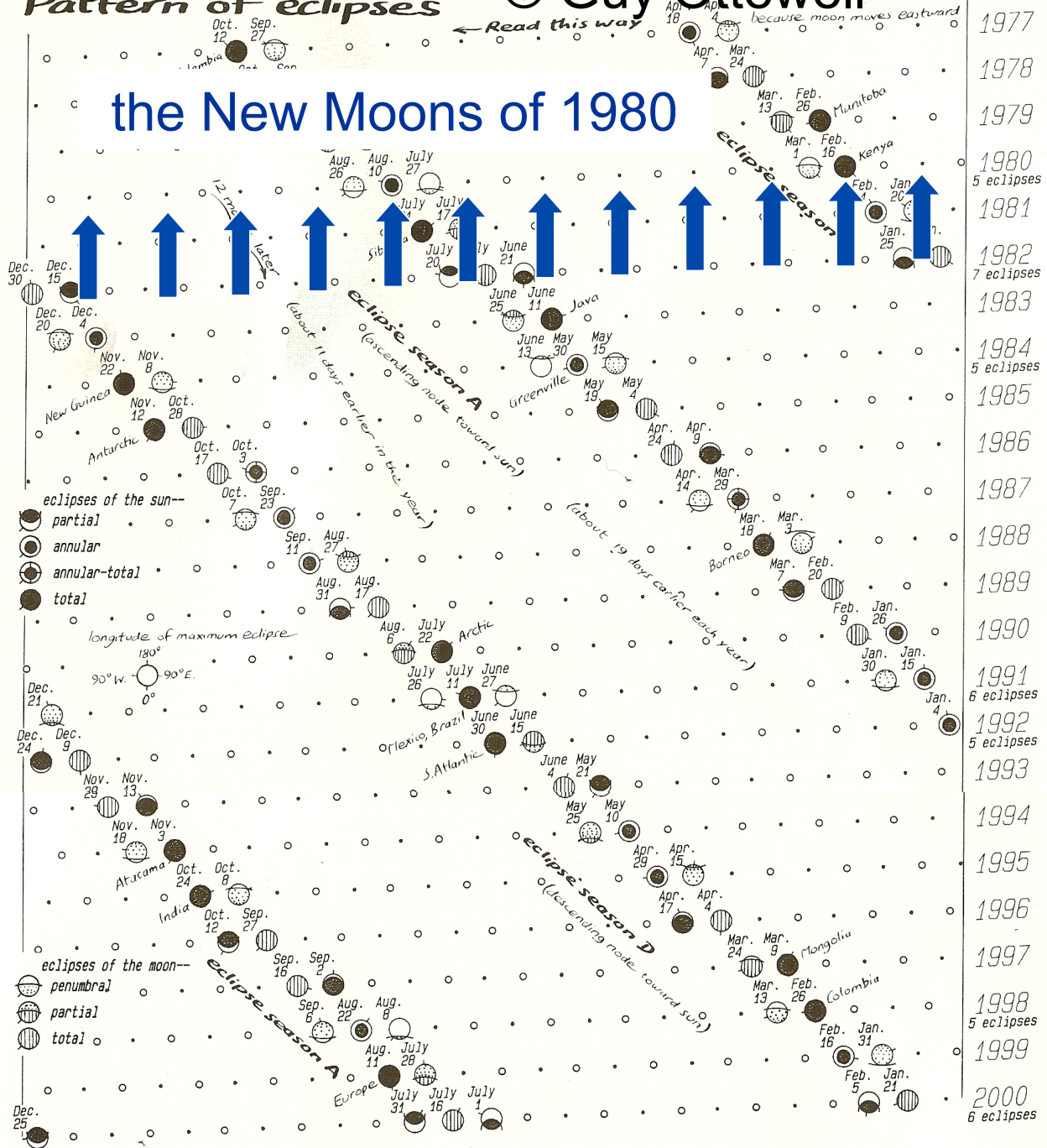
The Moon is in its full phase *significantly before or after* it crosses the ecliptic and encounters at least some part of the penumbra but never touches the umbra.



# Pattern of eclipses

© Guy Ottewell

## the New Moons of 1980



6585.221 days  
242 nodical months  
239 anomalistic months  
229 synodic months  
38 eclipse-seasons  
18 years  
2 half-saros  
one saros





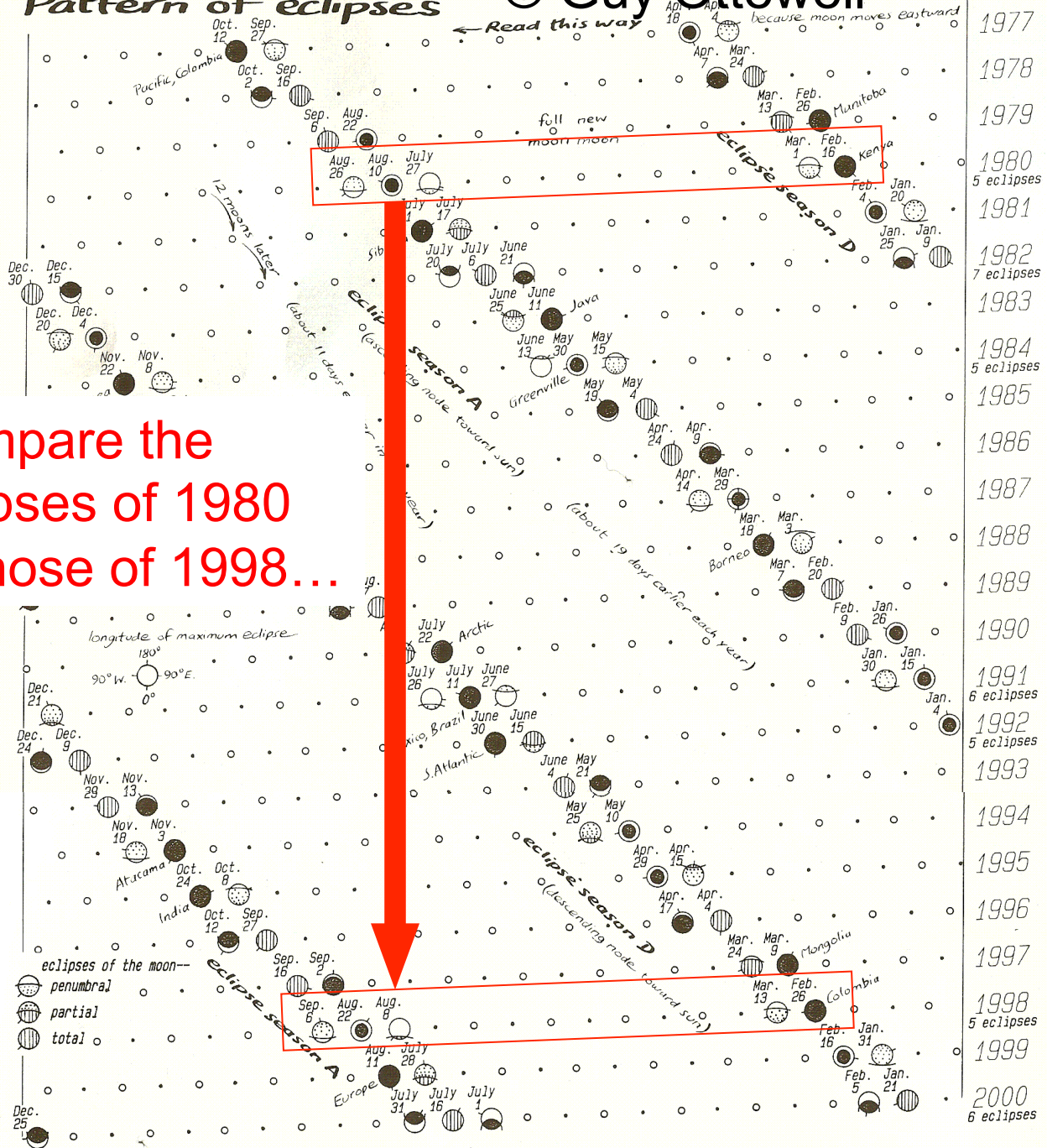






© Guy Ottewell

# Pattern of eclipses

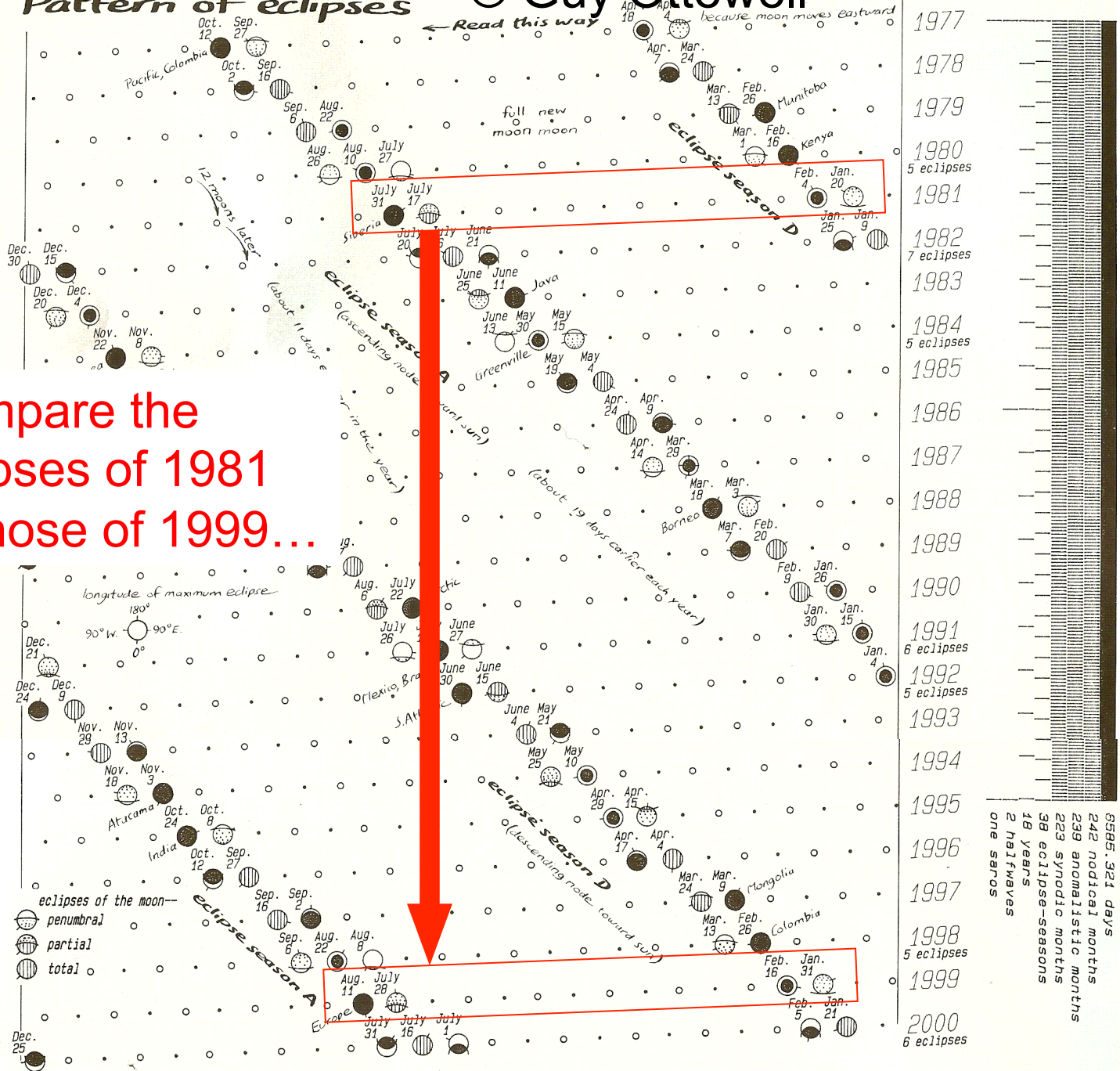


Compare the eclipses of 1980 to those of 1998...

6585.221 days  
 242 nodical months  
 239 anomalistic months  
 229 synodic months  
 38 eclipse-seasons  
 18 years  
 2 half-years  
 one saros



# Pattern of eclipses © Guy Ottewell



Compare the eclipses of 1981 to those of 1999...

← Read this way

because moon moves eastward

full new moon moon

eclipse season D

eclipse season A

(ascending node)

(about 11 days earlier in the year)

(about 19 days earlier each year)

eclipse season B

(descending node toward sun)

eclipse season A

eclipses of the moon—  
 penumbral  
 partial  
 total

longitude of maximum eclipse  
 180°  
 90°W 90°E  
 0°

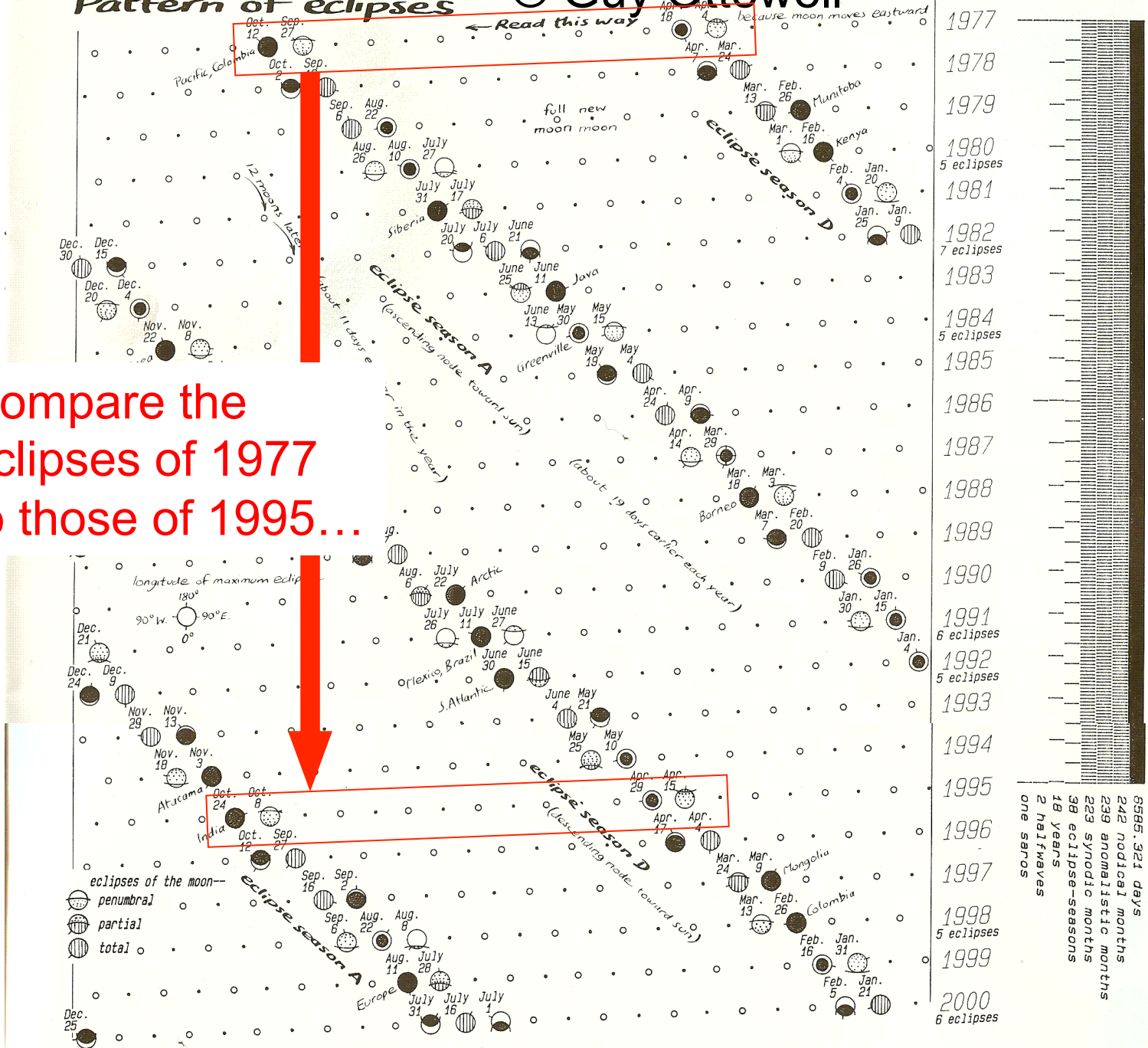
Dec. 25

1977  
 1978  
 1979  
 1980  
 5 eclipses  
 1981  
 1982  
 7 eclipses  
 1983  
 1984  
 5 eclipses  
 1985  
 1986  
 1987  
 1988  
 1989  
 1990  
 1991  
 6 eclipses  
 1992  
 5 eclipses  
 1993  
 1994  
 1995  
 1996  
 1997  
 1998  
 5 eclipses  
 1999  
 2000  
 6 eclipses





# Pattern of eclipses © Guy Ottewell



Compare the eclipses of 1977 to those of 1995...

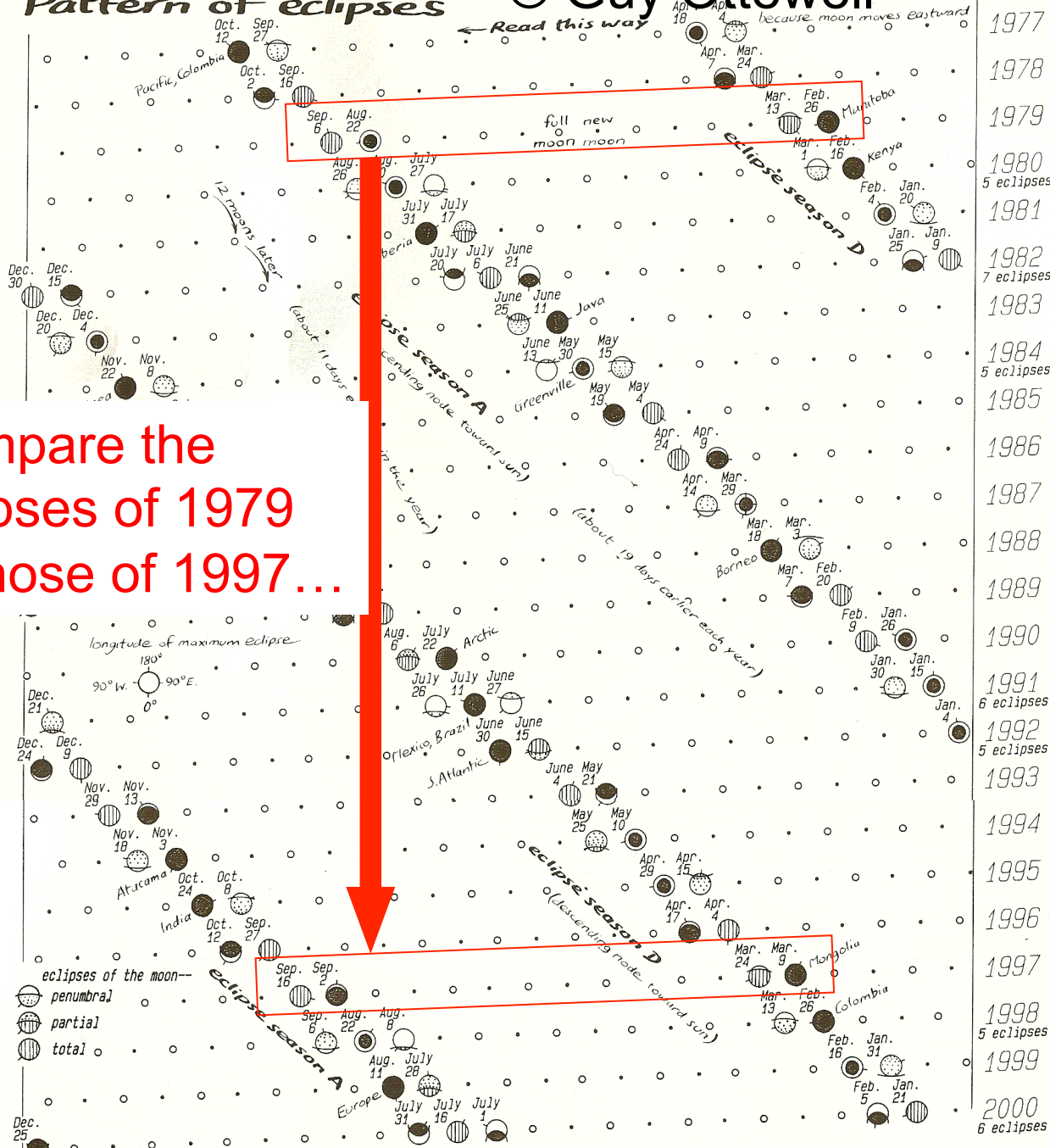
6585.221 days  
 242 nodical months  
 239 anomalistic months  
 229 synodic months  
 38 eclipse-seasons  
 18 years  
 2 half-years  
 one saros





© Guy Ottewell

# Pattern of eclipses



Compare the eclipses of 1979 to those of 1997...

# the Saros

- Every 18 years 11 days a pattern of eclipses is repeated. This is called “the Saros”.
- The solar eclipses every 18 years are of the same type, similar duration, and follow similar paths across Earth. However, the path will be about  $120^\circ$  farther west in longitude.
- A similar 18 year pattern exists for the lunar eclipses as well.
- Each saros series is numbered and studied by astronomers.



# Total and Annular Solar Eclipse Paths: 2001 – 2020

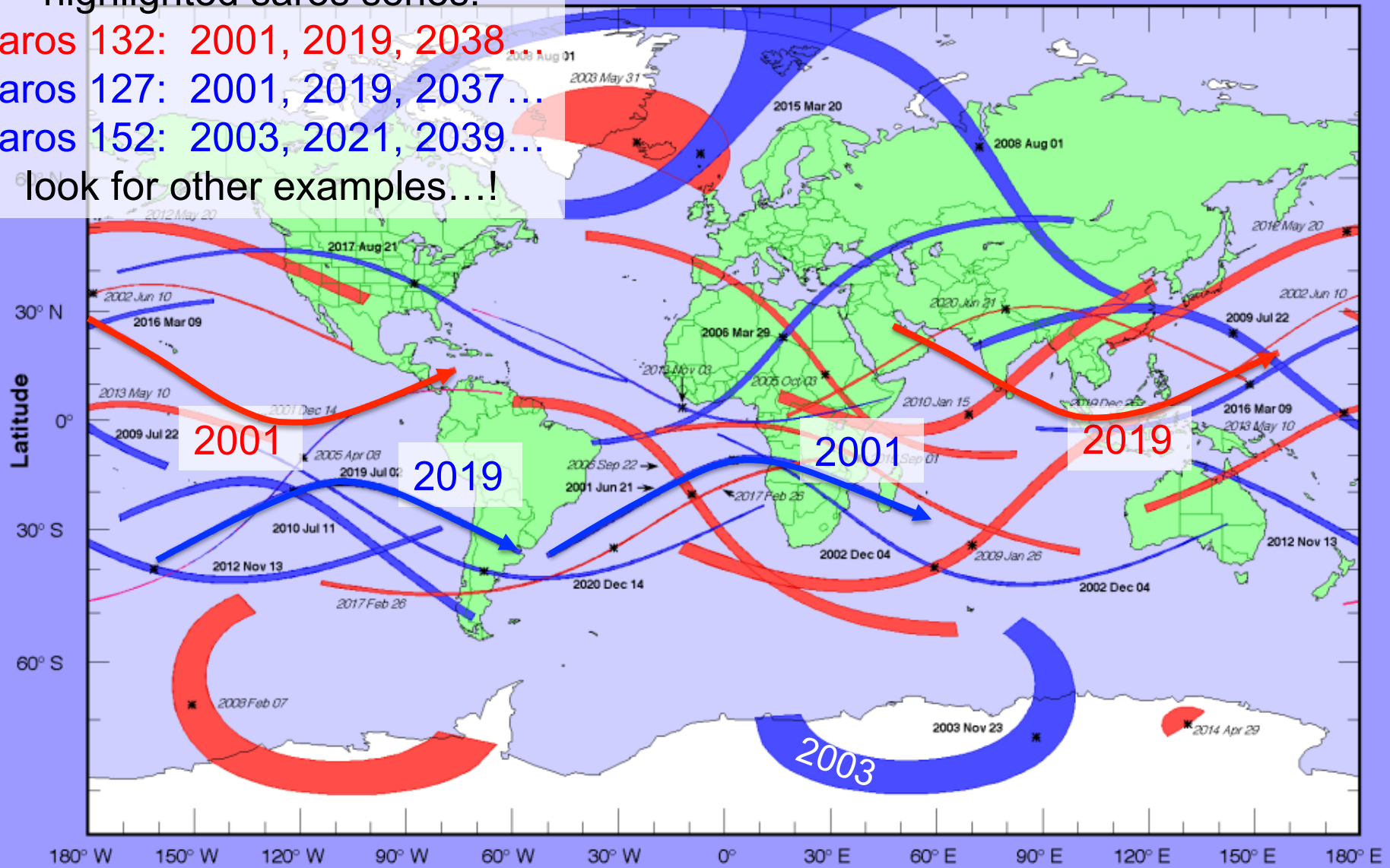
highlighted saros series:

saros 132: 2001, 2019, 2038

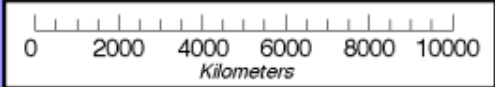
saros 127: 2001, 2019, 2037...

saros 152: 2003, 2021, 2039...

look for other examples...!



<span style="color: blue;">■</span>	<b>Total Eclipse</b>
<span style="color: red;">■</span>	<b>Annular Eclipse</b>
<span style="color: magenta;">■</span>	<b>Hybrid Eclipse</b>



# Total and Annular Solar Eclipse Paths: 2021 – 2040

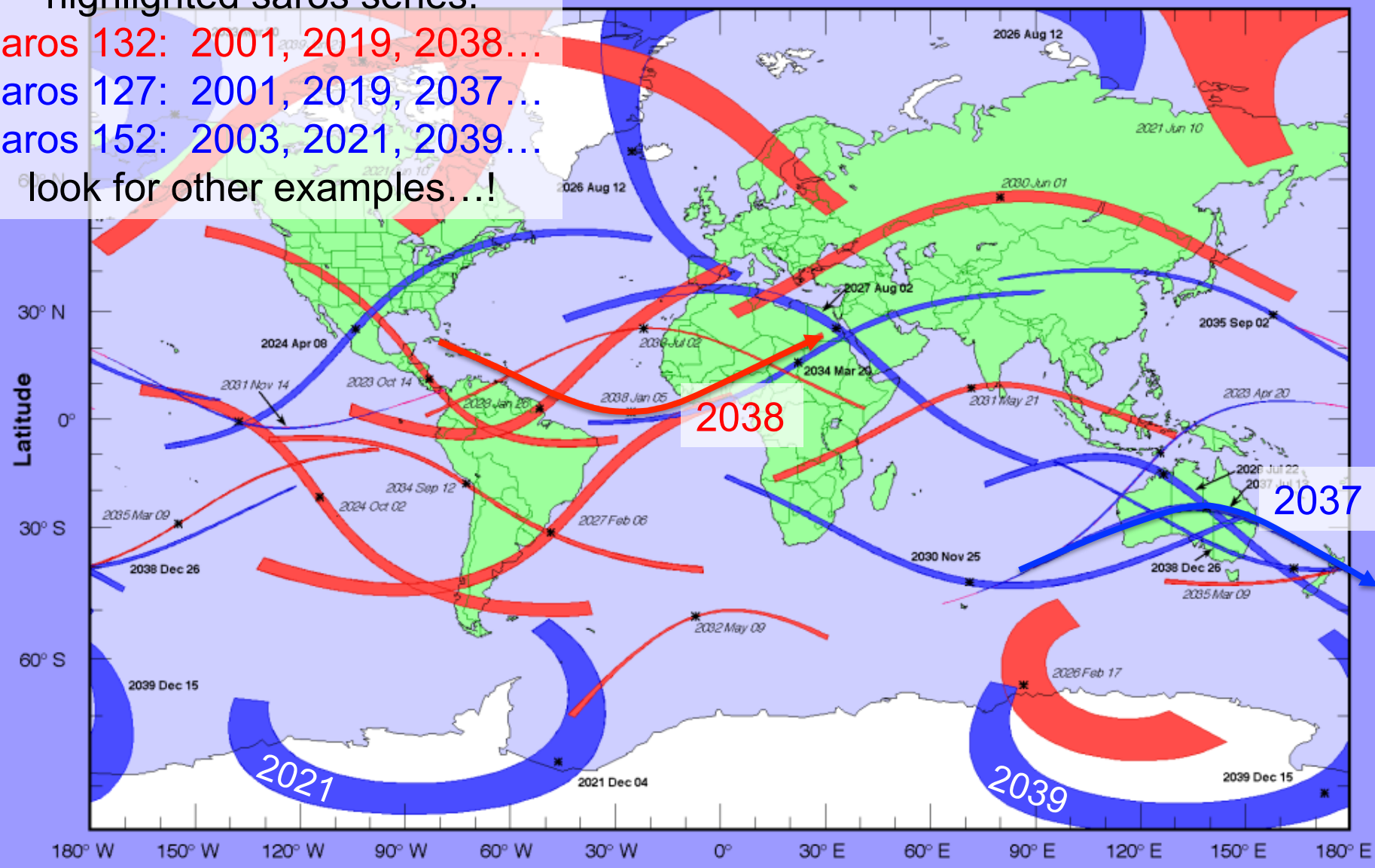
highlighted saros series:

saros 132: 2001, 2019, 2038...

saros 127: 2001, 2019, 2037...

saros 152: 2003, 2021, 2039...

look for other examples...!



	<b>Total Eclipse</b>
	<b>Annular Eclipse</b>
	<b>Hybrid Eclipse</b>

