Periodic Motion

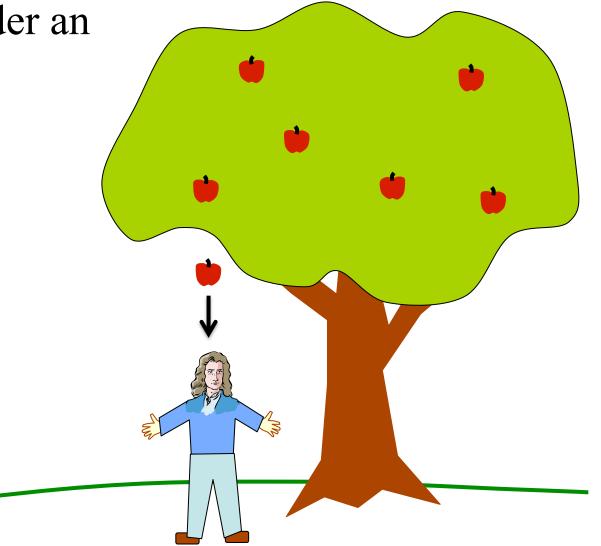
- I. Circular Motion
 - kinematics & centripetal acceleration
 - dynamics & centripetal force centrifugal force
- II. Universal Gravitation - Newton's "4th" Law
 - force fields & orbits
- III. Simple Harmonic Motionpendulums, springs, etc

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1	Define and calculate period and frequency.	
2	Apply the concepts of position, distance, displacement, speed, velocity, acceleration, and force to circular motion.	
3	State and correctly apply the relation between speed, radius, and period for uniform circular motion.	
4	State and correctly apply the relation between speed, radius, and centripetal acceleration for uniform circular motion.	1 – 14
5	Distinguish and explain the concepts of centripetal vs. centrifugal force.	15 –16
6	State and apply Newton's Law of Universal Gravitation.	17-28
7	Combine equations of circular motion and gravitation to solve problems involving orbital motion.	29-37
8	State and apply the relation between length, period, and g for a pendulum.	38-40
9	Solve problems involving application of Hooke's Law to the periodic motion of a mass attached to a spring. Also state and apply the relation between mass, period, and spring constant.	41 – 43

Let's put ourselves in Newton's shoes...



Once upon a time, Isaac Newton was under an apple tree...



<u>2nd Law:</u>

In order to accelerate down there must be force down.

In order to fall equally, gravity on an apple with twice the mass must be twice as great.

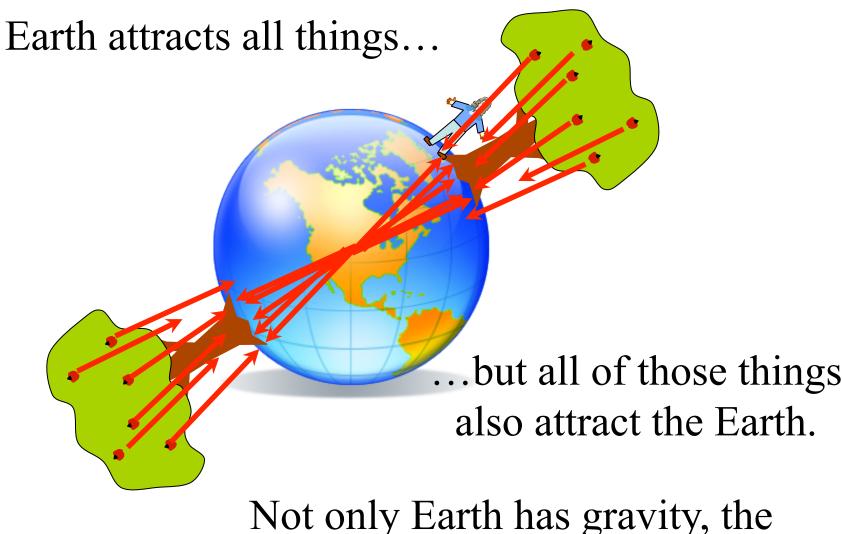
The force of gravity is proportional to mass! $(F_{\rm G} = mg)$

<u>3rd Law:</u>

All forces are interactions, so it is Earth that pulls the apple.

If Earth pulls apple, then apple pulls Earth!

It stands to reason the pull of the apple on the Earth depends on the mass of the Earth!



apple also has gravity, and any other thing with mass has gravity!



1st Law:

Moon would fly off into space *if* there were *no* force.



Moon's curved path shows that it is accelerating

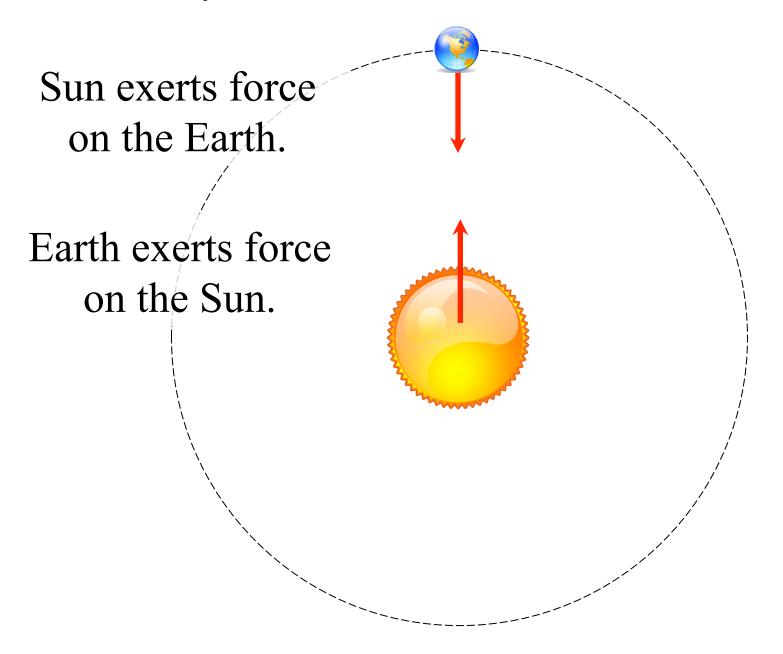
2nd Law:

Moon' s acceleration shows there *must* be a force *toward* the Earth. <u>3rd Law:</u>

Earth exerts force on Moon. Moon exerts equal and opposite force on Earth.



Similarly, Earth must accelerate toward the Sun!

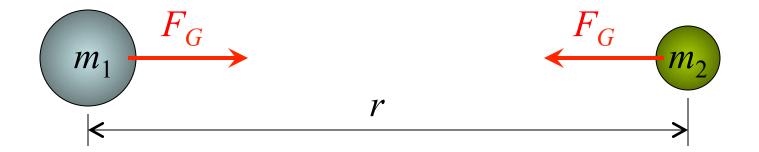


Newton's Law of Universal Gravitation:



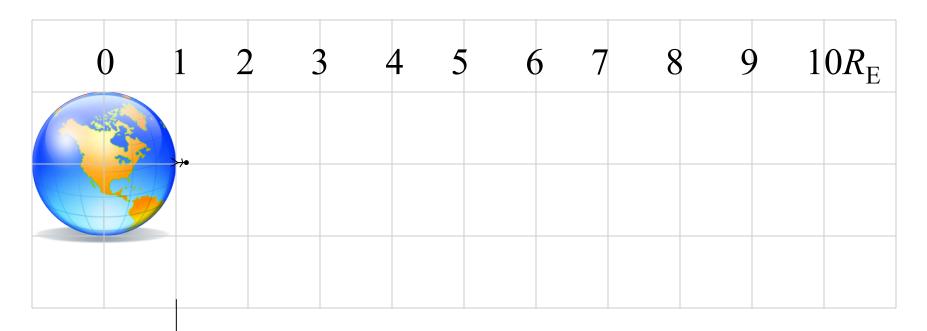
Every object in the universe attracts every other object in the universe.

Newton's Law of Universal Gravitation:



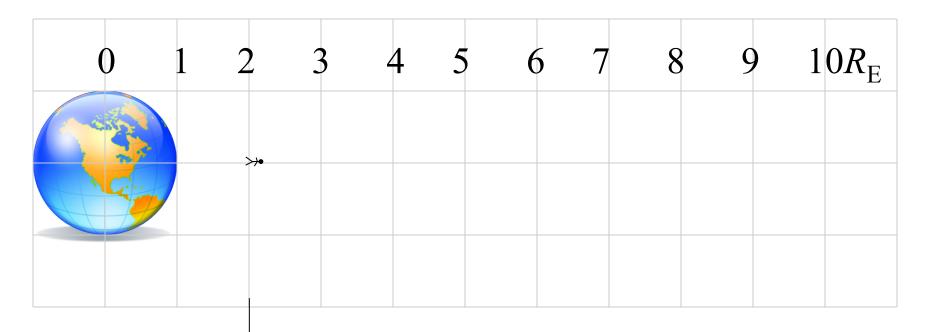
$$F_G = G \frac{m_1 m_2}{r^2}$$

where: m = mass r = distance between centers G = universal gravitational constant: $6.674 \times 10^{-11} \text{ m}^3/\text{kg s}^2$

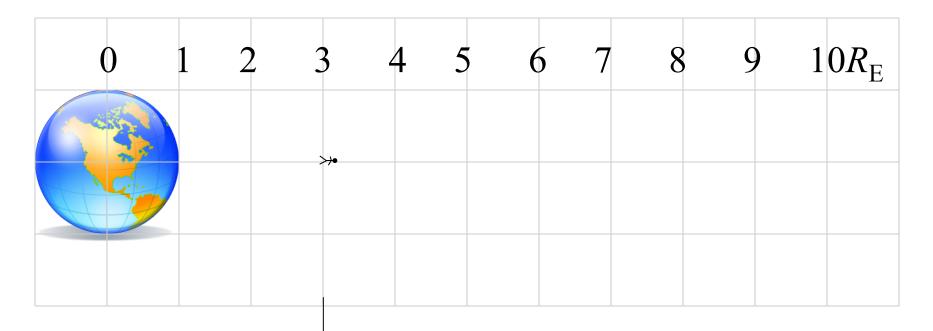


720 N

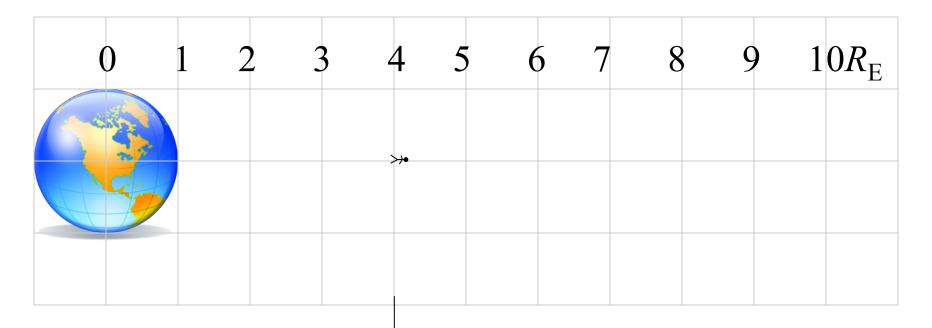
Consider a person that weighs 720 N on the surface of the Earth, who is then transported away to greater and greater distances...



180 NAt twice the distance the $\frac{720}{2^2} = 180$ person's mass is unchanged,but his weight is one fourthits original value.

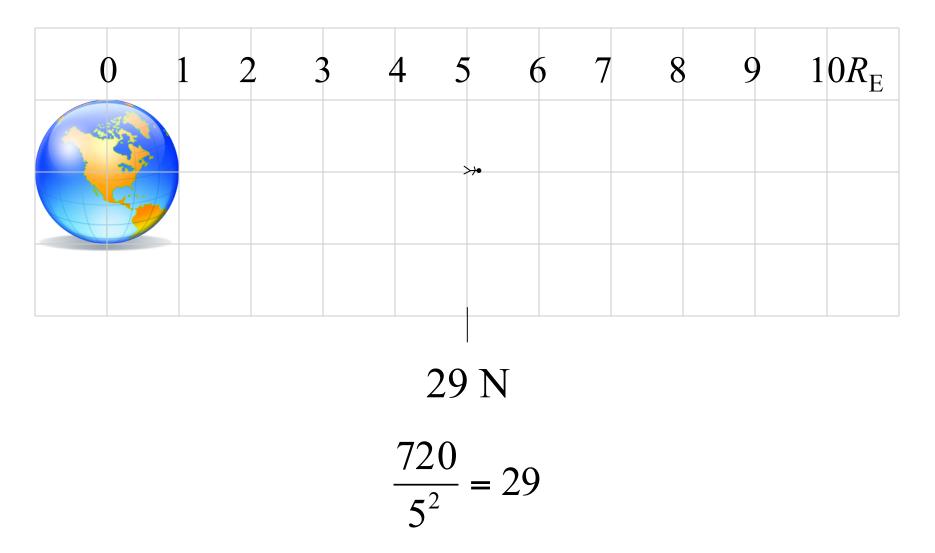


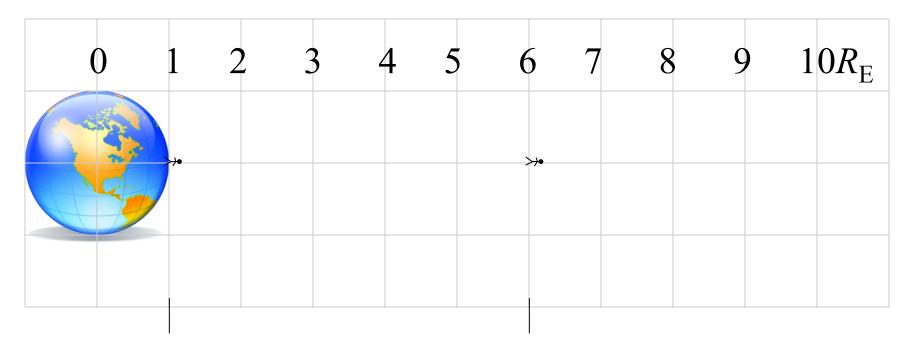
80 N $\frac{720}{3^2} = 80$



45 N

$$\frac{720}{4^2} = 45$$

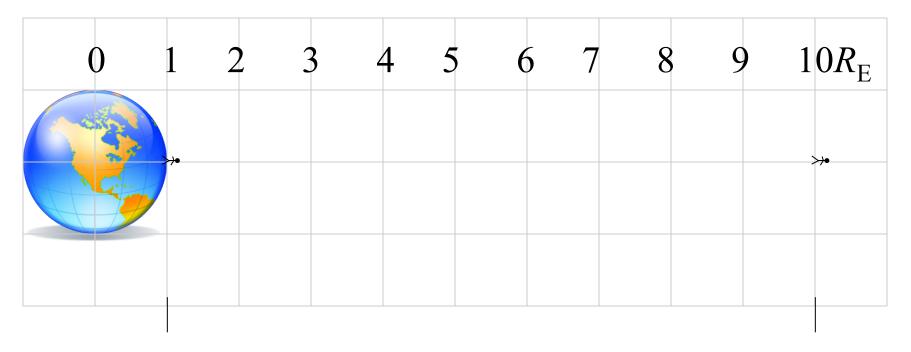




720 N

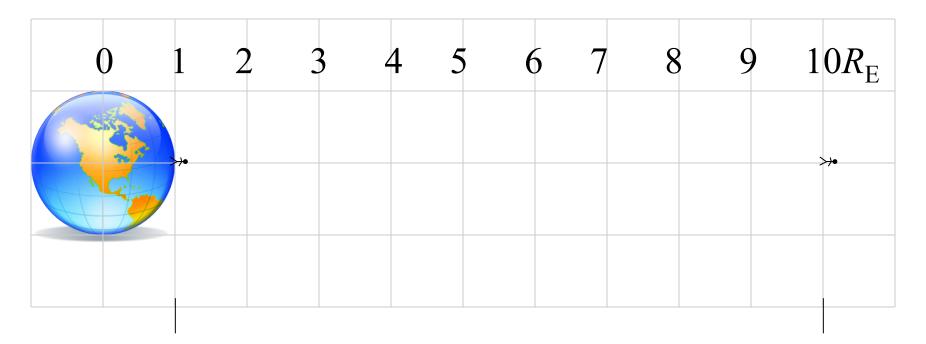
20 N

$$\frac{720}{6^2} = 20$$



720 N

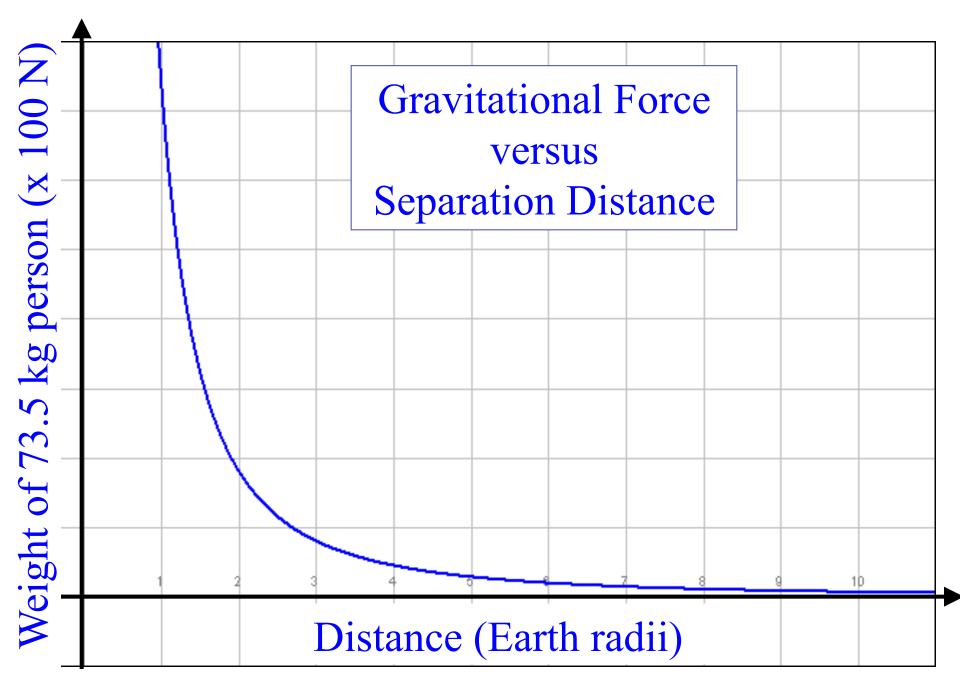
?



720 N

7.2 N

$$\frac{720}{10^2} = 7.2$$



Henry Cavendish determined G in 1798

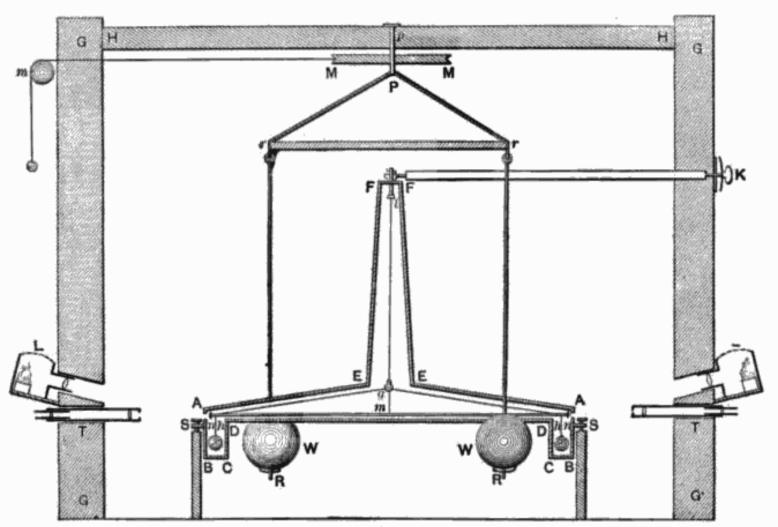
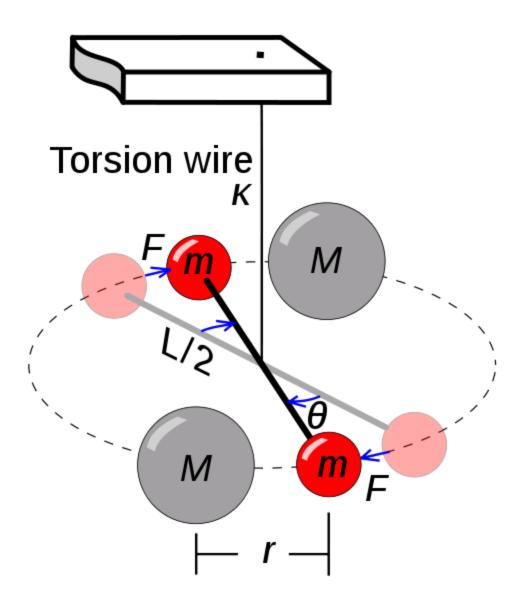


Fig. 1



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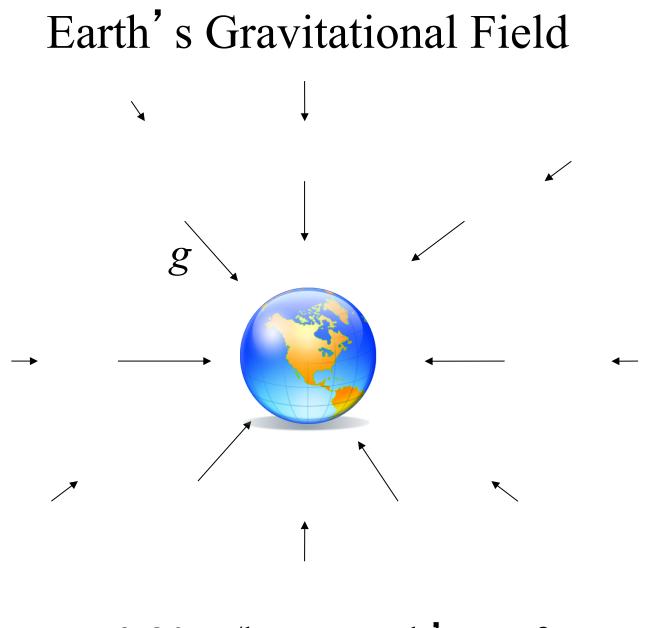
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Force Fields

A field is a region where a particular force has influence. There are many types of fields in physics: gravitational fields, magnetic fields, electric fields, etc.

Field strength is defined as the amount of force per unit of affected property.

A field has direction the same as the force. A field can be visualized as a series of vectors or lines that point in the direction of the force.



g = 9.80 N/kg at Earth' s surface

Gravitational Field



$$g = G \frac{M}{r^2}$$

where: M = mass of the field's *source* r = distance between centers G = universal gravitational constant: $6.674 \times 10^{-11} \text{ m}^3/\text{kg s}^2$

Interactive Satellite Viewer



GPS BIIA-26 (PRN 10) Satellite ID: 23953 3.903 Velocity (km/s): Velocity (mi/s): 2.425 Latitude (°): 44.117 Longitude(°): -175.494 Height (km): 19993.809 Height (mi): 12423.573 The orbit of a satellite is controlled by gravity. The centripetal acceleration is g.





A circular orbit for an object under the sole influence of gravity satisfies three equations:

Satell Veloci Veloci Latitu Longit Heigh Heigh $2\pi r$ \boldsymbol{a}

GPS BIIA-26 (PRN 10)

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